

Stars, Galaxies & the Universe Announcements

- **Reading Quiz #11** – in class today
- **HW#10 in ICON** – due Friday (11/12) by 5 pm
– available now in ICON
- **Exam #3 next Wednesday**
– study materials available by Friday (11/12) of this week!
- **Final Exam will be cumulative; Thursday 16 Dec @7:30 am in VAN LR 1; 150 points – 50 questions @3 pts each! We will have a review session sometime during Finals Week.**

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Stars, Galaxies & the Universe Lecture Outline

Expansion of the Universe

- (1) Hubble's Law and its Implications

Quasars & Active Galaxies

- (1) What is a quasar?
- (2) Powering of active galaxies
- (3) Images and effect of active galaxies

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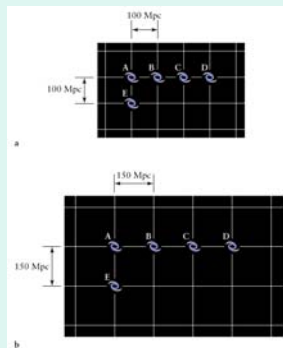
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- Overall, Hubble's Law tells you that the further away the galaxy, the faster the galaxy is receding – **universe is expanding!**

- Although universe is expanding, it looks the same from ANY galaxy – **we are not at the center!**

- The galaxies themselves are not getting bigger, expanding, but the **SPACE** between them is expanding



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Original wavelength

(a) A wave drawn on a rubber band ...

Stretched (redshifted) wavelength

(b) ... increases in wavelength as the rubber band is stretched.

→ photons get *redshifted* as they travel toward us from distant parts of the universe

10 Nov 2010 this is because the SPACE has actually expanded: photons stretched!

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There is much **heated debate** on the value of “Hubble’s Constant” H_0 – depends on the distance method used

- SN Type I method of finding distance – low values < 60 km/s/Mpc
- Tully-Fisher relation of finding distance – high values > 80 km/s/Mpc

Dr. Allan Sandage
 $H_0 = 58 \text{ km/s/Mpc}$

Dr. Wendy Freedman
 $H_0 = 74 \text{ km/s/Mpc}$

$H = 70$

Distance (Mpc)

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Hubble’s Constant (H_0) gives you AGE of universe!!

→ think about the REVERSE: galaxies rushing toward each other

Dr. Allan Sandage
 $H_0 = 58 \text{ km/s/Mpc}$

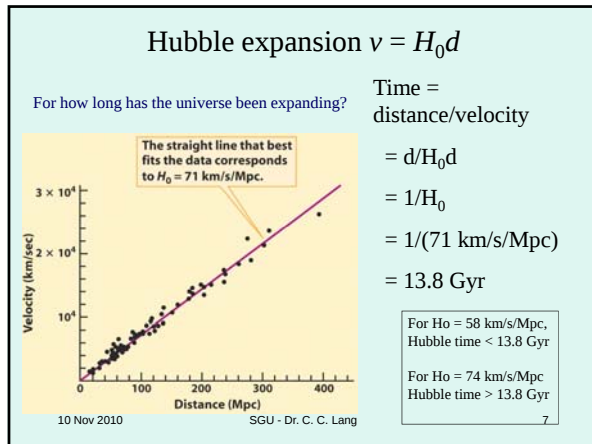
Dr. Wendy Freedman
 $H_0 = 74 \text{ km/s/Mpc}$

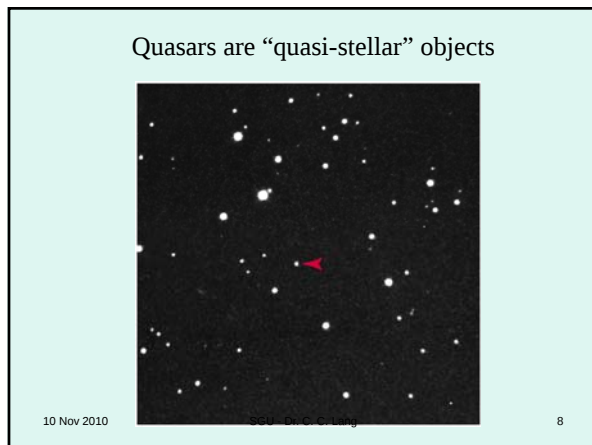
$H = 70$

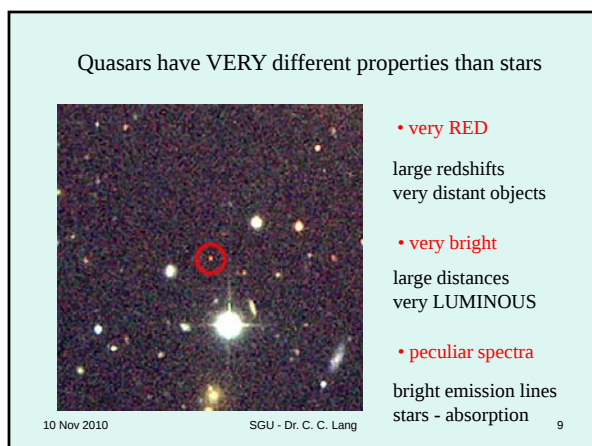
Distance (Mpc)

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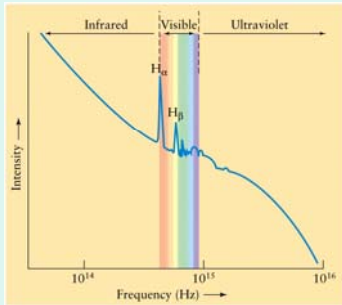
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Emission spectrum: detecting glowing gas

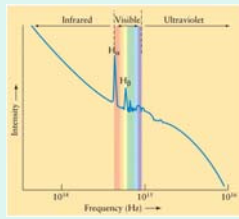
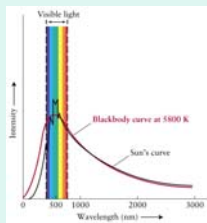


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Very strong lines at several wavelengths
No blackbody shape – emits strongly at many wavelengths

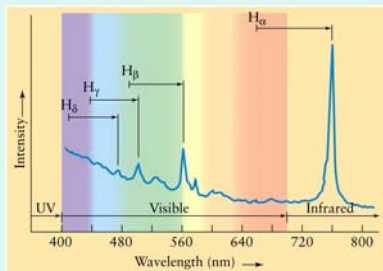


Wavelengths of the emission lines **REDSHIFTED**
by large amounts: what does this imply?

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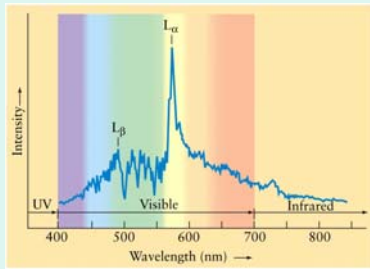
example: *moderately red shifted galaxy*

486 nm (blue-green) → 563 nm (yellow)

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example: *highly redshifted galaxy*

100 nm (ultraviolet) → 563 nm (yellow)

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Quasars are **VERY luminous** objects!

distance: large red shifts of spectral lines

flux: very large because we can easily detect them

$$\text{flux} = L / 4\pi d^2$$

LUMINOSITY: must be very large – intrinsic!

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examples of luminosities of quasars:

Luminosity of the Milky Way: $2.5 \times 10^{10} L_{\odot}$
(25 billion)

Luminosity of a typical quasar (3c273):
 $2.5 \times 10^{13} L_{\odot}$ (25 trillion)

Table 27-2 Galaxy and Quasar Luminosities	
Object	Luminosity (watts)
Sun	4×10^{26}
Milky Way Galaxy	10^{37}
Seyfert galaxies	$10^{36} - 10^{38}$
Radio galaxies	$10^{36} - 10^{38}$
Quasars	$10^{38} - 10^{42}$

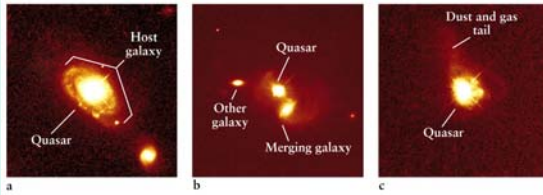
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Proof that quasars are associated with galaxies:

- can detect signatures of "host galaxies" around them
- see spiral arms, structure of the host galaxy



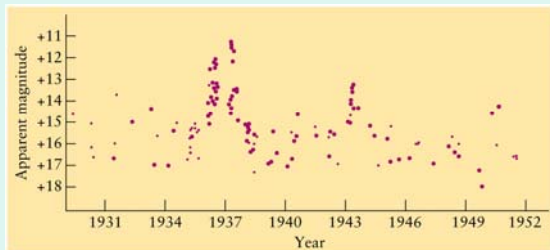
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Flickering on timescales of years

– *must be very small for a GALAXY*

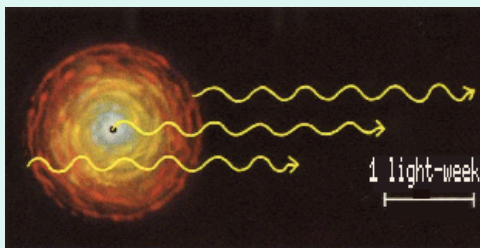


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The "flickering" is caused by light from the explosive event in a quasar leaving different parts of the galaxy and arriving at different times!

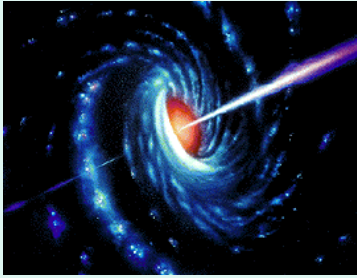


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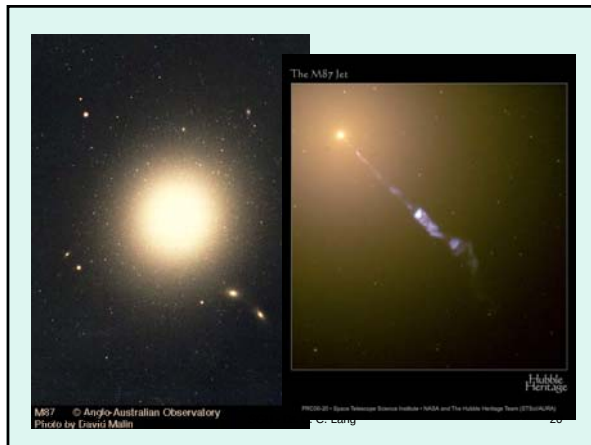
What can be causing the LARGE LUMINOSITIES
from such small objects?



Accretion of galactic gas onto the accretion
disk near a supermassive black hole!

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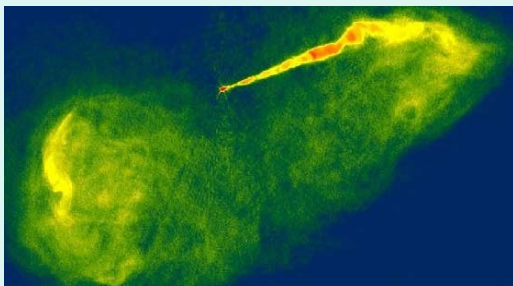
M87 © Anglo-Australian Observatory
Photo by Gerard Meert

THE M87 JET
NASA and The Hubble Heritage Team (STScI/AURA)

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12,000 light years across



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