

Stars, Galaxies & Universe

Lecture #4

- Electromagnetic radiation
 - Why we study it
 - What we can learn from it
 - Basic properties: speed, wavelength/frequency
 - The electromagnetic spectrum
 - The photoelectric effect
(energy = frequency)
 - The Doppler effect

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Basics of Electromagnetic Radiation (or “Light”)

- What we can learn from observing EM radiation from an astronomical object
 - *Temperature*
 - *Energy*
 - *Line-of-sight motion*
- Properties of EM Radiation
 - *How it interacts with matter*
 - *Wavelike properties (wavelength, frequency; Doppler Effect)*
 - *Particle-like properties (how light carries energy & how much: photoelectric effect, Wien’s Law)*

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Optical telescopes in the Canary Islands, Spain

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Arecibo Radio Telescope, Puerto Rico

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VLA radio telescope, New Mexico

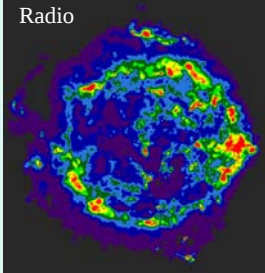


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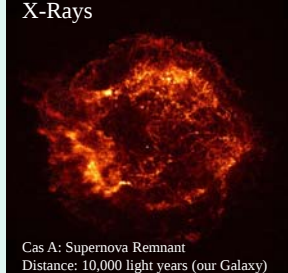
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Radio



X-Rays



Cas A: Supernova Remnant
Distance: 10,000 light years (our Galaxy)

- Astronomers collect “light” from astronomical objects
 - “light” usually refers to *VISIBLE* electromagnetic radiation
 - radio and x-rays are two other types of *EM* radiation
- We derive physical properties of objects: energetics, size, velocities

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Physical properties of EM radiation

- EM radiation travels at a constant speed IN A VACUUM

$$c = 299,792.458 \text{ km/s (186,282.397 miles/s)}$$

$$c \text{ (practical)} = 3 \times 10^5 \text{ km/s} = 3 \times 10^8 \text{ m/s}$$

- EM radiation travels slower in air, water, glass, etc.
→ space is nearly a vacuum, so mostly we use the value c



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Light from the

Travel time

- “light” from astronomical objects tells us about the physical properties when the light was emitted
- in this way, we are looking ‘back in time’
- the further away, the ‘older’ the view

Earth to the moon	1.28 seconds
Sun to Earth	8.5 minutes
Sun to Mercury	3 minutes
Sun to Venus	6 minutes
Sun to Mars	12.5 minutes
Sun to Jupiter	43 minutes
Sun to Saturn	1 hour
Sun to Uranus	2.6 hours
Sun to Neptune	4 hours
Sun to Pluto	5.4 hours
Sun to the nearest star	4.3 years
Sun to the most distant stars	14 billion years

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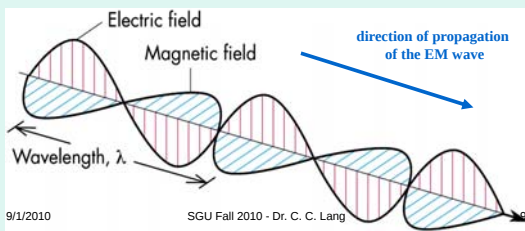
• EM radiation has wave-like properties

- waves travel through space
- waves are a disturbance in space

• What is doing the “waving”?

An EM wave is both an electric and magnetic disturbance

- EM wave emitted when charged particle (i.e. electron) accelerates
- EM waves create “fields” in space as they move



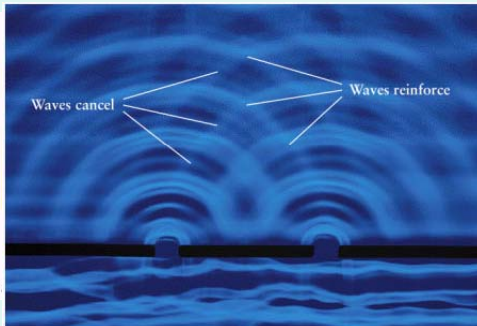
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Thomas Young's Double Slit Experiment: 1801

- showed that EM radiation had wave-like properties



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How does "light" (or EM radiation) interact with matter?



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Emit – electric current in lightbulb
Absorb – hand absorbing light from bulb
Transmit/Pass through - window
Reflect/Scatter – off snow, dust

- EM waves are characterized by their wavelengths (λ) or their frequencies (ν)

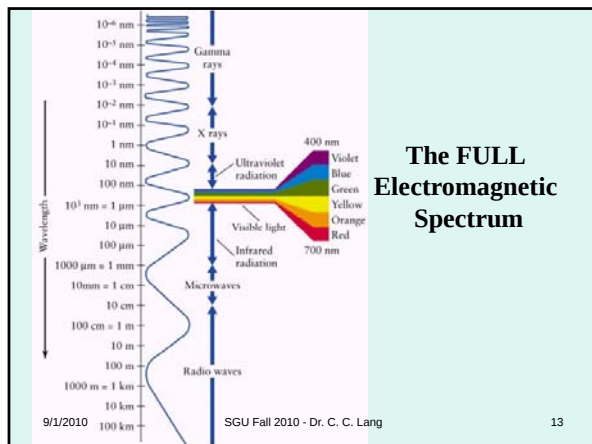
A prism, or "spectrometer", splits visible light 'white light' up into its different components



Much more sophisticated spectrometers are a main tool of astronomers

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• Which one has the **longest wavelength**?

Microwaves

Visible light

X-rays

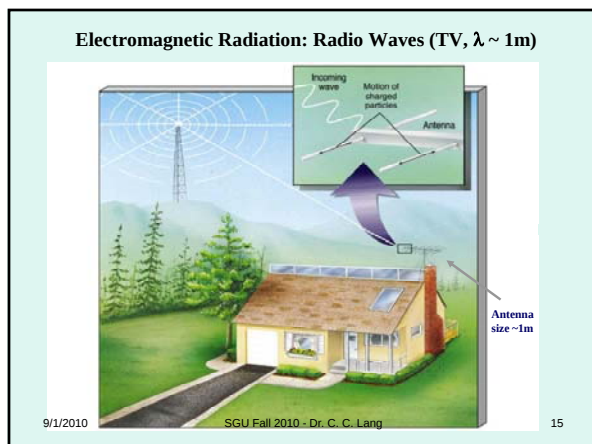
Ultraviolet radiation

Gamma-rays

Radio Waves

Infrared radiation

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Speed of light: relation between λ and ν

• speed of light = $c = \lambda \times \nu$

• $c = \text{constant} = 3 \times 10^8 \text{ m/s} = 3 \times 10^5 \text{ km/s}$

• λ has units of length (m)

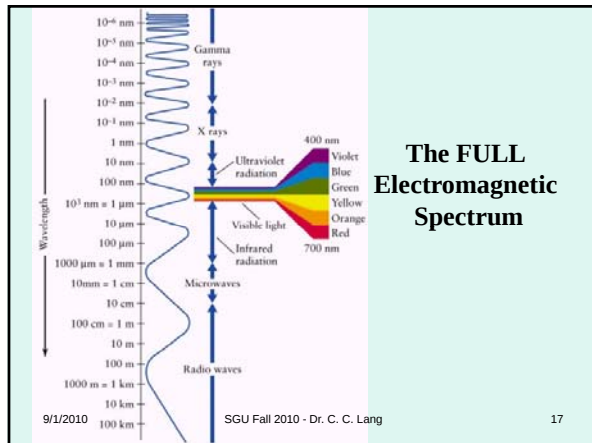
• ν has units of 1/s (Hz)

Example: what is the frequency of EM radiation with a wavelength (λ) of 50 micro-meters (μm)?

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• Which one carries the **most energy**?

Microwaves

Visible light

X-rays

Ultraviolet radiation

Gamma-rays

Radio Waves

Infrared radiation

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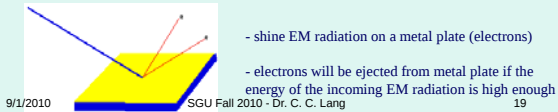
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EM radiation also has particle-like properties

- study of wave/particle duality: quantum mechanics

- A “photon” is another way of referring to EM radiation
 - used for describing its particle-like properties
 - can think of a photon like a packet of EM radiation
- The photoelectric effect demonstrates particle properties
 - “How we know EM radiation carries energy”
 - experiments being done in late 1800’s, Einstein 1905



- shine EM radiation on a metal plate (electrons)

- electrons will be ejected from metal plate if the energy of the incoming EM radiation is high enough

Photoelectric Effect:

- Outcome #1: if you increase **intensity** of light beam you get **more of the same electrons**
- Outcome #2: if you increase **frequency** of light beam you get **higher energy electrons**

→ Photons carry energy depending on their frequency (or wavelength)

$$E = h\nu = hc/\lambda$$

h is a physical constant, ν is frequency, E energy;
 c is speed of light and λ is wavelength

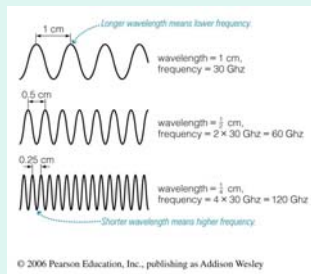
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E is proportional to ν : If ν increases, E increases

Do X-rays or radio waves carry more energy?

X-rays have a higher frequency (shorter wavelength)

and therefore X-rays carry **MORE ENERGY** than radio waves



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The Doppler Effect: wave-like property of EM radiation

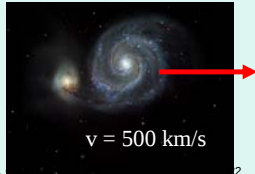
The Doppler effect is an apparent *change in wavelength (or frequency) of an EM wave*.

This change occurs *because of motion of the object emitting the EM radiation*, i.e. a star, a galaxy



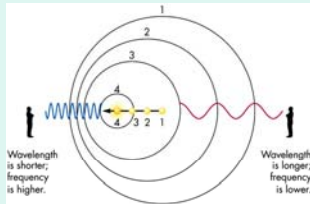
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The Doppler Effect: wave property



If object is moving **AWAY** from YOU, the wavelength appears **LONGER** – “**REDSHIFTED**”

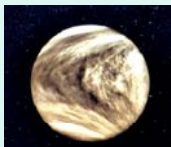
If object is moving **TOWARD** YOU, the wavelength appears **SHORTER** – “**BLUESHIFTED**”

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Doppler Shift examples – used throughout Astronomy!



Venus' rotation – discovered by the Doppler Shift of reflected radar waves



Doppler shift of sunlight
Sun's surface motions – entire Sun vibrates

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