

Midterm 1 Review:I. Magnetohydrodynamics (MHD)A. MHD Approximation

1. Strong Collisions: $\lambda_m \ll L, \tau \gg \nu_{ei}$
 2. Non-relativistic: $v_0^2/c^2 \ll 1$
 3. Magnetized: $r_{Li} \ll L$
- } Quasi-neutrality
} $\sum_s n_s q_s \approx 0$

B. Derivation from Two-fluid Equations

1. Drop $\underline{\underline{\Pi}}$, viscous terms from $\underline{\underline{P}}$ tensor (Collisional)
2. Zenerpi pressure (Collisional)
3. Drop Displacement Current (Non-relativistic)
4. Drop Hall term in Ohm's Law (Magnetized $r_{Li} \gg L$)
5. Adiabatic Equation of State

C. MHD Equations

1. Continuity $\frac{\partial \rho}{\partial t} + \underline{U} \cdot \nabla \rho = -\rho \nabla \cdot \underline{U}$

2. Momentum $\rho \frac{\partial \underline{U}}{\partial t} + \rho \underline{U} \cdot \nabla \underline{U} = -\nabla \left(p + \frac{B^2}{2\mu_0} \right) + \frac{(\underline{B} \cdot \nabla) \underline{B}}{\mu_0}$

3. Induction $\frac{\partial \underline{B}}{\partial t} = \nabla \times (\underline{U} \times \underline{B}) + \frac{\mu_0}{4\pi} \nabla^2 \underline{B}$

4. Adiabatic Equation of State $\frac{\partial p}{\partial t} + \underline{U} \cdot \nabla p = -\gamma p \nabla \cdot \underline{U}$

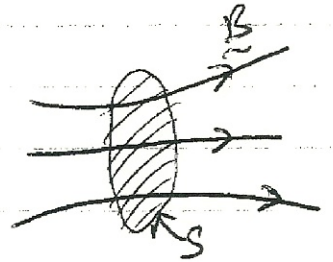
D. Frozen-In Flux

1. Magnetic Reynolds Number: $Re_M = \frac{\mu_0 L V_0}{\eta}$

2. Limits: a) $Re_M \gg 1 \Rightarrow$ Ideal MHD
 b) $Re_M \ll 1 \Rightarrow$ Diffusive

3. Frozen-In Flux Theorem: (Ideal Limit)

a. Magnetic Flux $\Phi_B = \int_S \vec{B} \cdot d\vec{A}$



b. Magnetic field lines are frozen to the flow

4. Clebsch Coordinates:

a. $\vec{B} = \nabla \alpha \times \nabla \beta$

b. $\alpha(x) = \text{const}$
 $\beta(x) = \text{const}$ } Defines field lines $\vec{B}(x)$

5. Applications of Frozen-In Flux: $\Phi_B = \text{constant}$ as dA changes!

E. MHD Linear Dispersion Relation

1. Know the steps to compute the linear dispersion relation!

a. Linearize the equations: $\rho = \rho_0 + \epsilon \rho_1(x)$, $\underline{U} = \underline{U}_0 + \epsilon \underline{U}_1(x)$,
uniform, steady ecc.

b. Fourier transform (plane-wave solution)

$$e^{i(\underline{k} \cdot \underline{x} - \omega t)} \Rightarrow \nabla \Rightarrow i \underline{k}$$

$$\frac{\partial}{\partial t} \Rightarrow -i\omega$$

I. Continued)

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c. Eliminate $p_{\perp}, \underline{B}_{\perp}, p_{\parallel}$ in terms of \underline{U}

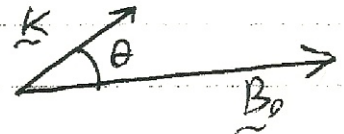
d. Collect to obtain matrix equation

$$\underline{M} \cdot \underline{U} = 0$$

e. Set Determinant $|\underline{M}| = 0 \Rightarrow$ yields $D(\omega, \underline{k}) = 0$

f. Solutions for ω (in terms of \underline{k} and other parameters) yield the linear wave modes!

F. MHD Wave Modes



$$1. \underline{k} = k_{\perp} \hat{x} + k_{\parallel} \hat{z} = k \sin \theta \hat{x} + k \cos \theta \hat{z}, \quad \underline{B} = B_0 \hat{z}$$

2. General Dispersion Relation

$$\underbrace{(\omega^2 - k_{\parallel}^2 v_A^2)}_{\text{Alfvén Modes, incompressible}} \left[\underbrace{\omega^4 - \omega^2 k^2 (c_s^2 + v_A^2) + k_{\parallel}^2 k^2 c_s^2 v_A^2}_{\text{Fast and Slow Magnetoacoustic Waves, compressible}} \right] = 0$$

Alfvén Modes,
incompressible

Fast and Slow Magnetoacoustic Waves,
compressible

$$\omega = \pm k_{\parallel} v_A$$

$$\frac{\omega^2}{k^2} = \frac{1}{2}(c_s^2 + v_A^2) \pm \frac{1}{2} \sqrt{(c_s^2 + v_A^2)^2 - 4c_s^2 v_A^2 \cos^2 \theta}$$

3. Sound speed

$$c_s^2 = \frac{\gamma p_0}{\rho_0}$$

Alfvén Speed

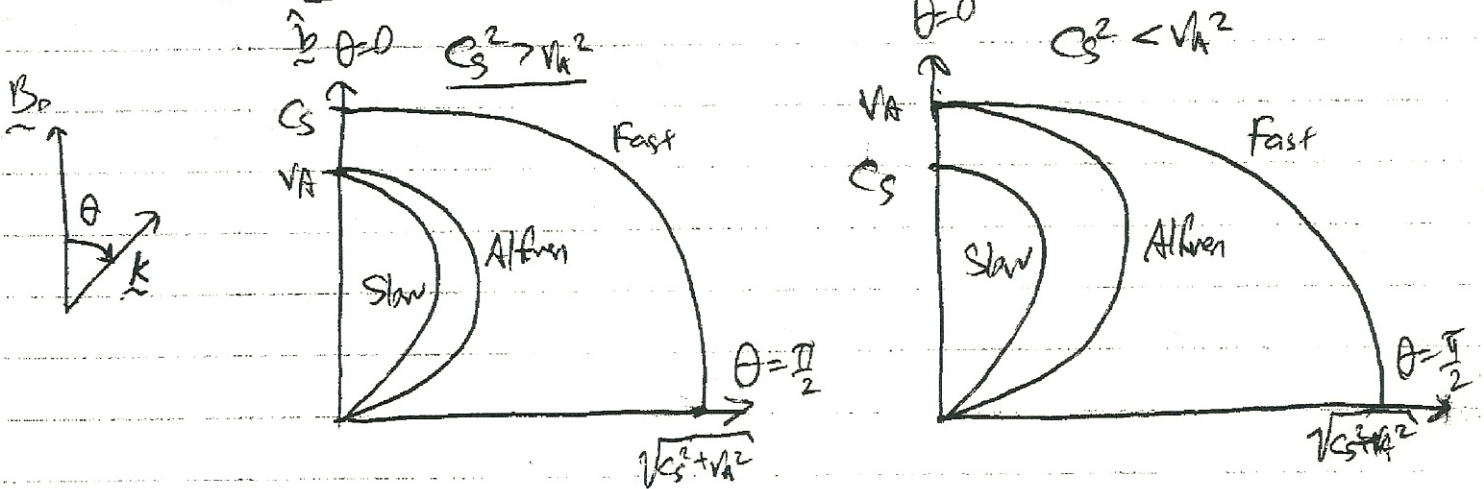
$$v_A^2 = \frac{B_0^2}{\mu_0 \rho_0}$$

4. a. Fast: Thermal & Magnetic Pressure add
Slow: Thermal & Magnetic Pressure oppose

Z.F. Continued)

Hawes (4)

5. Polar Plots: $v_p = \frac{\omega}{k} = \text{phase speed}$



G. Conserved Energy

$$1. E = \int d^3x \left(\frac{1}{2} \rho U^2 + \frac{p}{\gamma - 1} + \frac{B^2}{2\mu_0} \right)$$

H. Eigenfunctions of MHD Waves

$$1. \begin{pmatrix} (-) & 0 & (-) \\ 0 & (-) & 0 \\ (-) & 0 & (-) \end{pmatrix} \begin{pmatrix} U_x \\ U_y \\ U_z \end{pmatrix} = 0$$

2. Take $U_y = 0$, $U_x = U_0$

3. Solve for all other quantities in terms of U_x using the linearized, Fourier-transformed equations

$\Rightarrow p_1, \underline{u}_1, \underline{B}_1, p_1$ in terms of $U_x = U_0!$

4. Plug in ω for particular wave mode.

II. MHD Equilibrium

1. Force-Balanced Equilibrium: $\nabla p = \underline{j} \times \underline{B}$

2. Pfaff's Thm: Torus is simple surface
such that $\underline{B} \cdot \nabla p = 0$.

3. Force-Free Equilibrium: $\underline{j} \times \underline{B} = 0$

$$a \Rightarrow \nabla \times \underline{B} = \alpha \underline{B} \quad \Rightarrow \underline{B} \cdot \nabla \alpha = 0$$

surfaces of constant α

4. Cylindrical Equilibria: $\nabla \left(p + \frac{B^2}{2\mu_0} \right) = \frac{(\underline{B} \cdot \nabla) \underline{B}}{\mu_0}$
 p_r p_B μ_0 Mag. constant
 a. $\underline{B} = B_\phi(r) \hat{\phi} + B_z(r) \hat{z}$ (Ex. Flux rope, RFP)

b. $\frac{d}{dr} \left(p + \frac{B_\phi^2}{2\mu_0} + \frac{B_z^2}{2\mu_0} \right) = -\frac{B_\phi^2}{\mu_0 r}$

c. $B_z = 0 \Rightarrow z$ -pinch

$B_\phi = 0 \Rightarrow$ Theta-pinch

5. Toroidal Equilibria: (Nac on Eqm)

a. Magnetic Flux Coordinates $(\psi(r, z), F(r))$

b. Grad-Shafranov Equation

\Rightarrow Magnetostatic Equilibria

III. Multiple Timescale Methods

A. General Approach

1. Two timescales: $t, \tau = \epsilon^2 t$
2. $\frac{d}{dt} = \frac{\partial}{\partial t} + \epsilon^2 \frac{\partial}{\partial \tau}$
3. Expand variables: $X = \cancel{X_0} + \epsilon X_1 + \epsilon^2 X_2 + \dots$
4. Substitute into equation and multiply out to get all terms
5. Collect terms order by order to obtain
 - $\mathcal{O}(1)$
 - $\mathcal{O}(\epsilon)$ equations
 - $\mathcal{O}(\epsilon^2)$
6. Solve Equations order by order for x_1, x_2, \dots
 \Rightarrow set $\epsilon = 1$!
7. NOTE! Every problem is different. I will try to walk you through any problems or give you (Krook, Homes)

III. (Continued)

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B. Polarization Drift

1. Slowly varying electric field, $\underline{E}(t)$ with constant $\underline{B} = B_0 \hat{z}$.

$$2. \underline{v} = v_L (\cos(\omega t + \phi) \hat{x} - \sin(\omega t + \phi) \hat{y}) + \underbrace{\frac{\underline{E}(t) \times \underline{B}}{B_0^2}}_{\substack{\mathcal{O}(E) \\ \underline{E} \times \underline{B} \text{ drift}}} + \underbrace{\frac{1}{\omega_c B_0} \frac{d\underline{E}}{dt}}_{\substack{\mathcal{O}(E^2) \\ \text{Polarization} \\ \text{Drift, } \underline{v}_p}}$$

$$3. \underline{v}_p = \frac{1}{\omega_c B} \frac{d\underline{E}}{dt} = \frac{m}{2B^2} \frac{d\underline{E}}{dt}$$

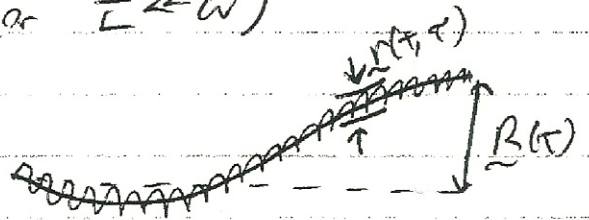
C. Ponderomotive Force

1. Particle motion in rapidly varying electromagnetic field

$$\underline{E}(\underline{x}, t, \tau) = \underline{E}_0(\underline{x}, \tau) \cos(\omega t - \underline{k} \cdot \underline{x})$$

2. Ordering: $\epsilon = \frac{v}{L\omega} \ll 1$ (or $\frac{v}{L} \ll \omega$)

3. $\underline{x} = \underbrace{\underline{R}(\tau)}_{\text{oscillation center}} + \epsilon \underline{\rho}(t, \tau)$



4. $\underline{F}_{\text{pond}} = m \frac{d\underline{u}}{dt} = \frac{-q^2}{4\pi\epsilon_0 \omega^2} \nabla |\underline{E}_0|^2$

5. $\underline{F}_{\text{pond}} = -\nabla \Phi_{\text{pond}} \quad \Phi_{\text{pond}} = \frac{q^2}{4\pi\epsilon_0 \omega^2} |\underline{E}_0|^2$

NOTE:
Drop $\cos(\omega t - \underline{k} \cdot \underline{x})$ in this part!

6. Energy Picture: $E_{\text{osc}} = \frac{1}{2} m u^2 + \Phi_{\text{pond}} = \text{constant!}$

IV. MHD Stability

A. Ideal MHD Equations, Cont

$$\text{Mom. } \rho \left[\frac{\partial \underline{U}}{\partial t} + \underline{U} \cdot \nabla \underline{U} \right] = -\nabla p + \frac{(\nabla \times \underline{B}) \times \underline{B}}{\mu_0}$$

$$\text{Ind } \frac{\partial \underline{B}}{\partial t} = \nabla \times (\underline{U} \times \underline{B})$$

$$\text{Press. } \frac{\partial p}{\partial t} + \underline{U} \cdot \nabla p = -\gamma p \nabla \cdot \underline{U}$$

$$2. E = \int d^3x \left[\frac{1}{2} \rho U^2 + \frac{p}{\gamma-1} + \frac{B^2}{2\mu_0} \right]$$

B. Linear Force Operator

1. Plasma Equilibrium gives $\rho_0(\underline{x})$, $\underline{B}_0(\underline{x})$, $p_0(\underline{x})$ (where $\underline{U}_0=0$)

2. For a plasma displacement $\underline{\xi}_1$

$$a. \rho_0 \frac{\partial^2 \underline{\xi}_1}{\partial t^2} = \underline{F}(\underline{\xi}_1)$$

$$b. \underline{F}(\underline{\xi}_1) = \nabla \left[\underline{\xi}_1 \cdot \nabla p_0 + \gamma p_0 \nabla \cdot \underline{\xi}_1 \right] + \frac{(\nabla \times \underline{B}_0) \times \left[\nabla \times (\underline{\xi}_1 \times \underline{B}_0) \right]}{\mu_0} \\ + \frac{(\nabla \times [\nabla \times (\underline{\xi}_1 \times \underline{B}_0)]) \times \underline{B}_0}{\mu_0}$$

C. Expand Energy in orders of $|\underline{\xi}_1|$

$$1. \mathcal{O}(|\underline{\xi}_1|^0) = \mathcal{O}(1)$$

$$E_0 = \int d^3x \left[\frac{p_0}{\gamma-1} + \frac{B_0^2}{2\mu_0} \right]$$

$$2. \mathcal{O}(|\underline{\xi}_1|^1)$$

$$E_1 = \int d^3x \left[\nabla p_0 \cdot \underline{\xi}_1 - \frac{(\nabla \times \underline{B}_0) \times \underline{B}_0 \cdot \underline{\xi}_1}{\mu_0} \right] = 0$$

$$3. \mathcal{O}(|\underline{\xi}_1|^2)$$

$$E_2 = \int d^3x \left[\frac{1}{2} \rho_0 \left| \frac{\partial \underline{\xi}_1}{\partial t} \right|^2 \right] + \delta W(\underline{\xi}_1, \underline{\xi}_1) \\ = \delta K + \delta W$$

IV, (Continued)
C. (Continued)

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4. Since $\underline{F}(\underline{\xi}_1)$ is self-adjoint

$$\int d^3x \underline{\xi} \cdot \underline{F}(\underline{\xi}) = \int d^3x \underline{\xi} \cdot \underline{F}(\underline{\xi})$$

$$\Rightarrow \delta W = \frac{1}{2} \int d^3x \underline{\xi} \cdot \underline{F}(\underline{\xi})$$

D. Normal Modes:

$$1. \rho_0 \frac{\delta^2 \underline{\xi}_1}{\delta t^2} = \underline{F}(\underline{\xi}_1) \Rightarrow -\rho_0 \omega_n^2 \underline{\xi}_n = \underline{F}(\underline{\xi}_n)$$

$$a. \omega_n^{2*} = \omega_n^2 \Rightarrow \omega_n^2 \text{ is real}$$

$$b. \int d^3x \underline{\xi}_n \cdot \underline{\xi}_m = \delta_{nm} \Rightarrow \text{orthonormal displacements!}$$

E. Energy Principle

$$1. E_2 = \delta K + \delta W$$

2. $\delta W \geq 0 \Rightarrow$ Necessary & Sufficient for stability

$\delta W < 0$ for any $\underline{\xi} \Rightarrow$ UNSTABLE

3. write $\delta W(\underline{\xi}, \underline{\xi})$ and minimize w.r.t. $\underline{\xi}$ (each component)

if $\delta W(\underline{\xi}_{\min}, \underline{\xi}_{\min}) < 0 \Rightarrow$ UNSTABLE

$$4. \delta W = \frac{1}{2} \int d^3x \left[\underbrace{|\nabla \times (\underline{\xi} \times \underline{b}_0)|^2}_{\text{Mag tension \& compression}} + \underbrace{\gamma \rho_0 |\nabla \cdot \underline{\xi}|^2}_{\text{Thermal compression}}$$

$$- \sum^* \int d^3x \left[\nabla \times (\underline{\xi} \times \underline{b}_0) \right] - \sum^* \int d^3x \left[\nabla(\underline{\xi} \cdot \nabla p_0) \right]$$

"Kink drive" $\leftarrow \rightarrow$ "Interchange drive"
potentially destabilizing