

Lecture #10 Kinetic Stability of a Plasma

Hawes ①

I. Kinetic Stability of a Plasma

A. Gardner's Theorem: A single-humped velocity distribution is always stable.

Proof: 1. Roots of dispersion relation $D(k, p) = 0$ give

real frequency and growth/decay rates of normal modes.

2. For a solution with $\text{Re}(p) > 0$, the plasma is unstable!

3. Proof by contradiction: Assume there are solutions with $\text{Re}(p) > 0$.

4. Since $\text{Re}(p) > 0$, then, for $k > 0$, we can take the linear dr integration along the $\text{Re}(V_r)$ axis

$$D(k, p) = 1 - \frac{\omega_p^2}{k^2} \int_{-\infty}^{\infty} dv_r \frac{\partial F_0 / \partial v_r}{V_r - \frac{\omega}{k} - \frac{i\gamma}{k}} = 0$$

where we have substituted $p = \gamma - i\omega$

5a. We can separate the integrand into Real & Imaginary parts

$$\frac{\partial F_0 / \partial v_r (V_r - \frac{\omega}{k} + \frac{i\gamma}{k})}{V_r - \frac{\omega}{k} - \frac{i\gamma}{k}} = \frac{\partial F_0 / \partial v_r (V_r - \frac{\omega}{k})}{(V_r - \frac{\omega}{k})^2 + \frac{\gamma^2}{k^2}} + i \frac{\gamma}{k} \frac{\partial F_0 / \partial v_r}{(V_r - \frac{\omega}{k})^2 + \frac{\gamma^2}{k^2}}$$

b. Thus

$$D_r(k, p) = 1 - \frac{\omega_p^2}{k^2} \int_{-\infty}^{\infty} dv_r \frac{\partial F_0 / \partial v_r (V_r - \frac{\omega}{k})}{(V_r - \frac{\omega}{k})^2 + (\frac{\gamma}{k})^2} = 0$$

$$D_i(k, p) = -\frac{\omega_p^2 \gamma}{k^2} \int_{-\infty}^{\infty} dv_r \frac{\partial F_0 / \partial v_r}{(V_r - \frac{\omega}{k})^2 + (\frac{\gamma}{k})^2} = 0$$

c. Both real & imaginary pieces must equal zero separately.

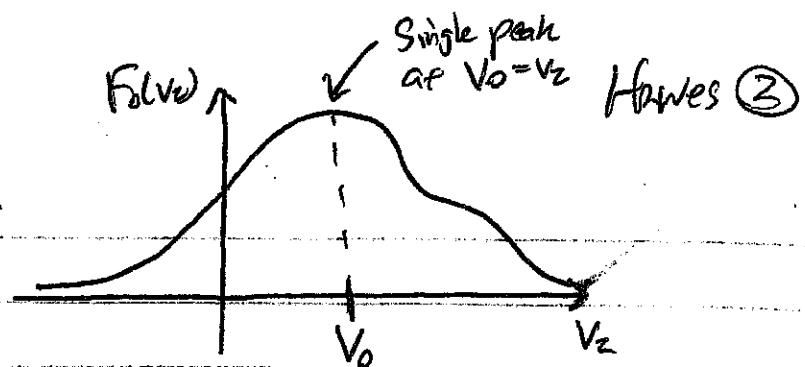
Thus any linear combination of D_r & D_i must equal zero.

$$\text{Take } D_r - \left(\frac{k V_0 - \omega}{\gamma}\right) D_i = 0$$

Lecture #10 (Continued)

I. A. (Continued)

6. Single-humped Verbiage distribution



$$7. \text{ Thus } D_r + \left(\frac{kV_b - \omega}{\gamma} \right) D_i = 1 - \frac{\alpha \nu^2}{k^2} \int_{-\infty}^{\infty} dv_r \frac{\frac{\partial F_0}{\partial v_r}}{(v_r - \frac{\omega}{k})^2 + (\frac{\gamma}{k})^2} \left[(v_r - \frac{\omega}{k}) - \left(\frac{kV_b - \omega}{\gamma} \right) \frac{\gamma}{k} \right]$$

$$\text{a. Pieces in brackets } [-] = v_r - \frac{\omega}{k} - V_b + \frac{\omega}{k} = v_r - V_b$$

b. Thus, we have

$$1 + \frac{\alpha \nu^2}{k^2} \int_{-\infty}^{\infty} dv_r \frac{\left(\frac{\partial F_0}{\partial v_r} \right) (V_b - v_r)}{(v_r - \frac{\omega}{k})^2 + (\frac{\gamma}{k})^2} = 0$$

8. NOTE: a. Denominator of integrand is positive definite

b. For single humped distribution

$$\left(\frac{\partial F_0}{\partial v_r} \right) (V_b - v_r) > 0 \quad \text{for } v_r > V_b \text{ and } v_r < V_b.$$

a. Thus, the integrand is positive definite, leading to a positive definite integral. Thus, the relation above can never be satisfied!

b. This contradiction means the original assumption, $\text{Re}(p) > 0$, is false.

c. Thus, the single-humped distribution is always stable.

QED,

B. Nyquist Criterion

i. If a distribution function with a single peak is stable, how do we test a multiply-humped distribution for stability?

2. Dispersion Relation: $D(k, p) = 0$ yields solutions.

If some k arises such that a solution has $\text{Re}(p) = \gamma > 0$, UNSTABLE!

Lesson #10 (Continued)

I. B. (Continued)

3. Complex p-plane

$$\text{Im}(p) = i\omega$$

UNSTABLE

Howes 3

a. Line $\gamma=0$ is boundary between stable and unstable.
⇒ Marginal Stability

b. Since $D(k, p)$ is a complex function of p , we can map the unstable ($\text{Re}(p) > 0$) half of the p -plane into complex $-D$ space.

c. This unstable half-plane is bounded by the $\gamma=0$ line from $\omega = -\infty$ to $\omega = +\infty$.

d. If the point $D=0$ falls within the mapping of unstable region, then an unstable solution exists.

4. Example: a. Cauchy Velocity Distributions $F_0(v_z) = \frac{C}{\pi} \frac{1}{(C^2 + v_z^2)}$

b. From Lecture #7, the dispersion relation is

$$D(k, p) = 1 + \frac{\omega p^2}{(p + ikC)^2}$$

c. Substituting $p = \gamma - i\omega$ and calculating D_r and D_i gives

$$D_r = 1 + \frac{\omega p^2 [(p + ikC)^2 - \omega^2]}{[(p + ikC)^2 + \omega^2]^2}$$

$$D_i = \frac{\omega p^2 2ikC\omega}{[(p + ikC)^2 + \omega^2]^2}$$

Mapping of $\gamma=0$

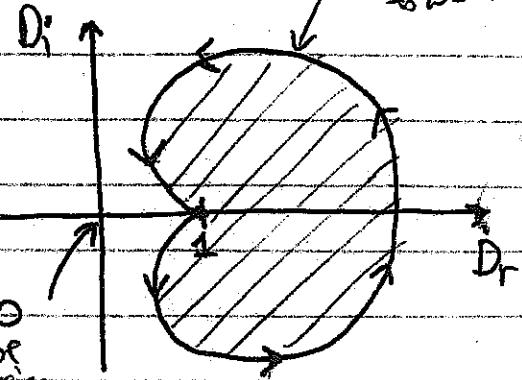
From $\omega = -\infty$
 $\rightarrow \omega = +\infty$

d. Setting $\gamma=0$ (Marginal Stability)
we obtain:

$$D_r = 1 + \frac{\omega p^2 (k^2 C^2 - \omega^2)}{[k^2 C^2 + \omega^2]^2}$$

$$D_i = 2 \frac{\omega p^2 ikC\omega}{[k^2 C^2 + \omega^2]^2}$$

$D=0$
is me
a solution



Lecture #10 (Continued)
Z. B. 4. (Continued)

Homework ④

e. Since $D=0$ is not within the mapping of the unstable $\text{Re}(p) > 0$ half-plane, the plasma is stable.

⇒ This result is consistent with Gardner's Theorem.

C. The Winding Theorem: If a closed contour C_p in the complex p -plane encloses n simple zeros of some mapping function $D(p)$, then the corresponding contour C_D in the complex D -plane must make n turns around the origin.

Proof:

1. From Residue theorem, number of turns, N_w , of contour C_D above the origin is

$$N_w = \frac{1}{2\pi i} \oint_{C_D} \frac{dD}{D}$$

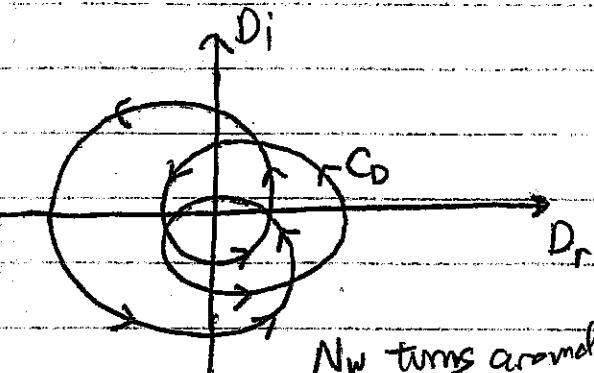
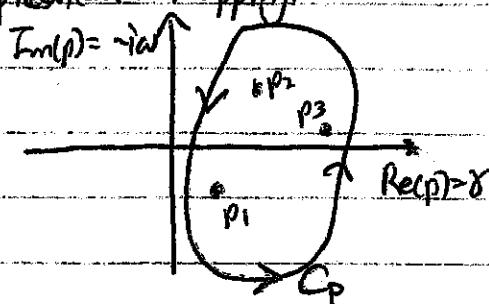
pde occurs at $D=0$.

Def: Winding Number

2. Changing variables to the p -plane ($dD = \frac{\partial D}{\partial p} dp$), we have

$$N_w = \frac{1}{2\pi i} \oint_{C_p} \frac{1}{D} \frac{\partial D}{\partial p} dp$$

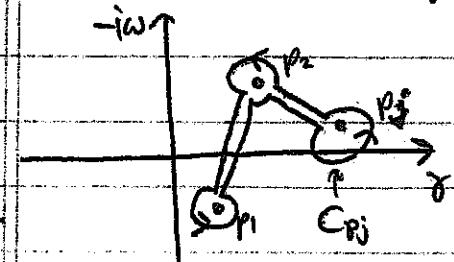
3. Representative Mapping



N_w turns around

$$D = D_r + iD_i = 0$$

f. Deform Contour C_p :



$$N_w = \sum_{j=1}^n \frac{1}{2\pi i} \int_{C_{pj}} \frac{1}{D} \frac{\partial D}{\partial p} dp$$

Lecture #10 (Continued)

I.C (Continued)

Hawes (5)

5.a Taylor Expand $D(p)$ above $p=p_j$

$$D(p) = D(p_j) + \frac{\partial D}{\partial p} \Big|_{p_j} (p-p_j) + \dots$$

Solution

b. We can also expand the function $g = \frac{\partial D}{\partial p}$ about $p=p_j$

$$g(p) = g(p_j) + (p-p_j) \frac{\partial g}{\partial p} \Big|_{p_j} + \dots$$

To lowest order, keep only $g(p_j) \Rightarrow \frac{\partial D}{\partial p} = \frac{\partial g}{\partial p} \Big|_{p_j}$

$$\text{c. Thus } \frac{1}{D} \frac{\partial D}{\partial p} = \frac{\left(\frac{\partial g}{\partial p}\right)_{p_j}}{(p-p_j) \left(\frac{\partial g}{\partial p}\right)_{p_j}} = \frac{1}{p-p_j}$$

$$\text{d. Thus, we obtain } N_W = \sum_{j=1}^n \frac{1}{2\pi i} \int_{C_{p_j}} \frac{dp}{p-p_j} = n \quad \checkmark \quad \text{QED.}$$

D. The Perron Condition

1. Let's Apply the Nyquist Criterion to a distribution function with an arbitrary number of humps.

Table where an expression for $D(k, p)$ valid for any distribution function.

b. Since γ is always small near $\gamma=0$, we can use the Plemelj Relation to evaluate $D_r(k, p)$ and $D_i(k, p)$. (We did this for Weak Gravitational Approximation, Lec #8, I.C.4.)

$$D_r(k, \omega) = 1 - \frac{\alpha p^2}{k^2} \int_{-\infty}^{\infty} P(z) dz \frac{\frac{\partial f_0}{\partial z}}{V_2 - \frac{\alpha z}{k}}$$

$$D_i(k, \omega) = -\pi \frac{k}{R} \frac{\alpha p^2}{k^2} \frac{\partial f_0}{\partial z} \Big|_{z=\frac{\alpha k}{R}}$$

Lecture #10 (Continued) I. D. (Continued)

Hawes ⑥

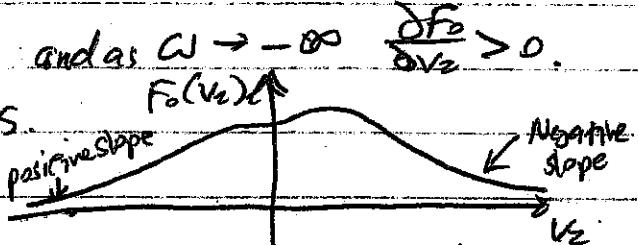
3. Shape of $\gamma=0$ curve in D-plane near $\omega = \pm\infty$

a. First, we'll assume $k > 0$.

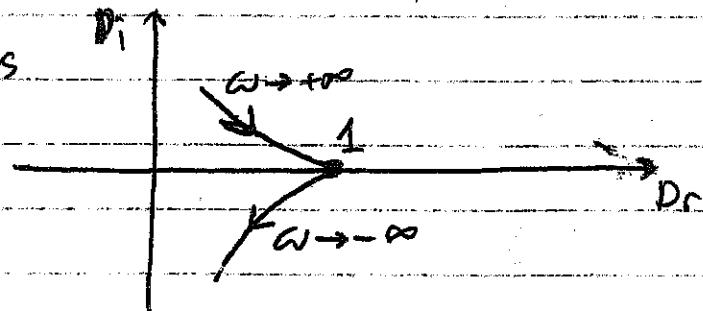
b. Note: $\frac{\partial F_0}{\partial v_2} \rightarrow 0$ as $\omega \rightarrow \infty$, so $\lim_{\omega \rightarrow \infty} D_r = 1$ $\lim_{\omega \rightarrow \infty} D_i = 0$.

c. Also, as $\omega \rightarrow \infty$ $\frac{\partial F_0}{\partial v_2} \leq 0$ and as $\omega \rightarrow -\infty$ $\frac{\partial F_0}{\partial v_2} > 0$.

Since $F_0(v_2) > 0$ always.



d. Thus



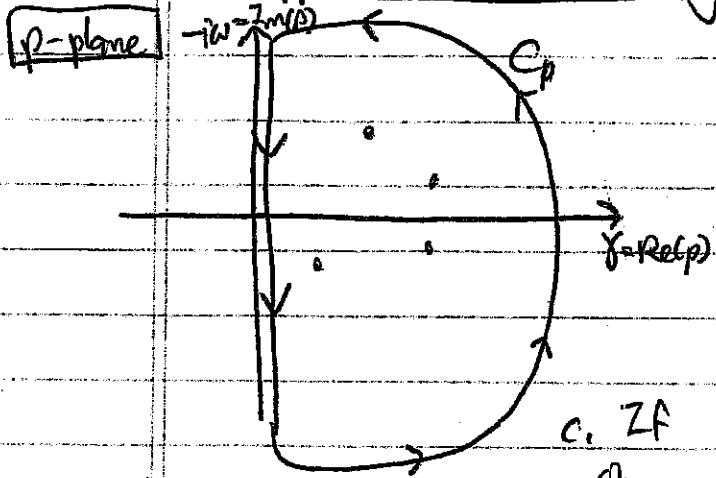
4. Crossing D_r -axis ($D_i = 0$)

a. Contour crosses D_r axis when $D_i = 0 \Rightarrow \frac{\partial F_0}{\partial v_2} = 0$

\Rightarrow Only crosses D_r where distribution function has zero slope.

b. For a smooth, continuous $F_0(v_2)$, there are always an odd number of crossings.

5. Application of the Winding Theorem



a. Take contour C_p downward along $\gamma=0$, closing at infinity in right half plane.
 \Rightarrow (Thus C_p encloses all instabilities)

b. Contour of $|p|= \infty$ maps to $D=1$.

Rest of contour maps to $\gamma=0$ curve.

c. If $\gamma=0$ curve in D-space winds around the origin ($D=0$) (ccw), then unstable root exists.

Lecture 11D (Continued)

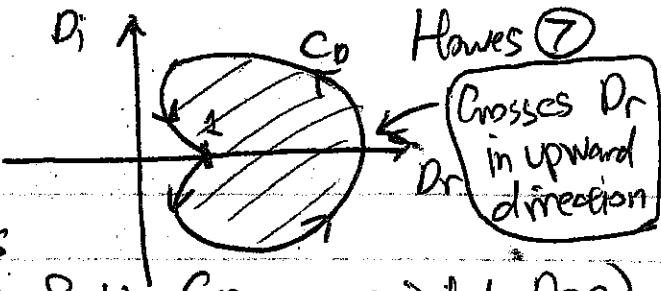
Z. D. (Continued)

G. For single-hump distribution:

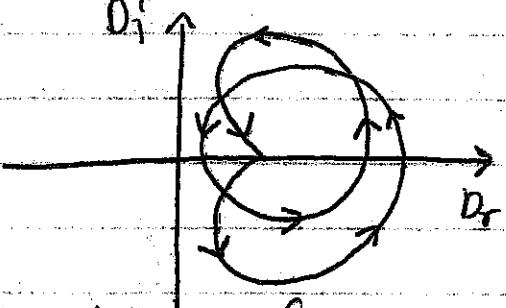
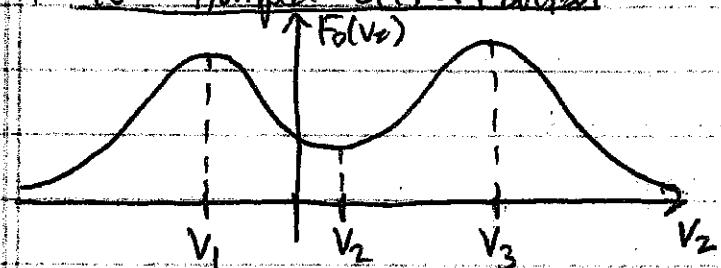
a. CCW contour C_D must cross

D_r -axis to the right of 1 \Rightarrow Stable (D_r does not include $D=0$).

\Rightarrow Once again, we have proven Gardner's Theorem.



7. Two-humped distributions



a. Cross D_r -axis three times: Upward at v_1 & v_3 (maxima)
Downward at v_2 (minimum)

b. At the points where $\frac{\partial F_0}{\partial v_z} = 0$, $D_i = 0$ and $v_z = \frac{c_0}{k} = V_j$,

Thus, the crossing of the D_r -axis occurs at

$$D_r = 1 - \frac{c_0^2}{k^2} \int_{-\infty}^{\infty} dv_z \frac{\frac{\partial F_0}{\partial v_z}}{v_z - V_j}$$

c. If $D_r < 0$ at a crossing, the plasma is unstable by the Nyquist criterion.

d. Since k may be arbitrarily small, we must have

$$\int_{-\infty}^{\infty} dv_z \frac{\frac{\partial F_0}{\partial v_z}}{v_z - V_j} > 0 \quad \text{for any } j \text{ to have instability.}$$

F. The Penrose Condition

a. Noting that $F_0(V_j) = \text{const}$, we may write

$$\int_{-\infty}^{\infty} dv_z \frac{\frac{\partial F_0}{\partial v_z}}{v_z - V_j} = \int_{-\infty}^{\infty} dv_z \frac{\frac{d}{dv_z} [F_0(v_z) - F_0(V_j)]}{v_z - V_j}$$

Lecture #10 (Continued)

I. D. 8. (Continued)

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b. Integrating by parts, we obtain

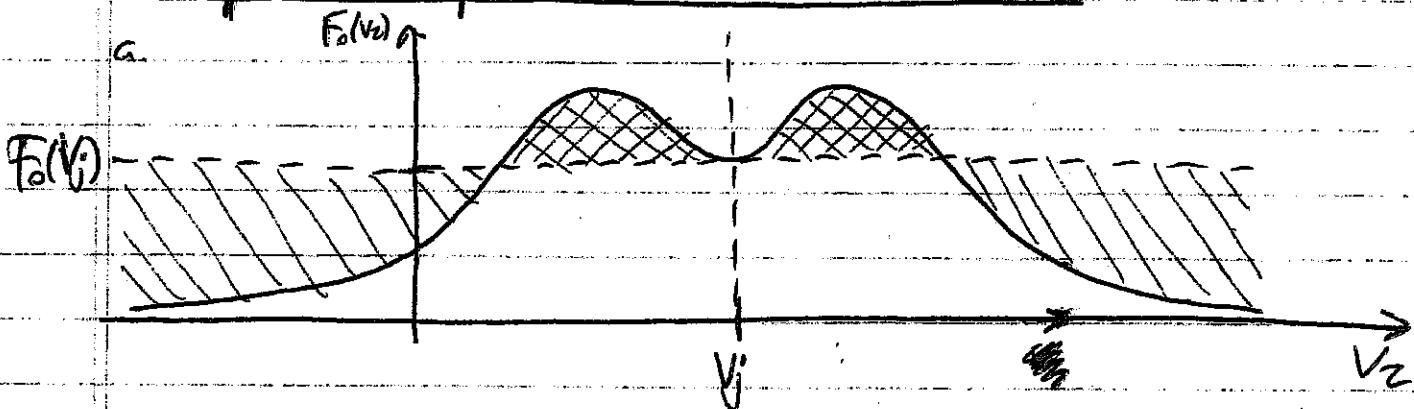
$$\int_{-\infty}^{\infty} dv_z \frac{F(v_z) - F_0(v_j)}{(v_z - v_j)^2} > 0$$

Penrose Condition
for Inseability

c. NOTE! We may drop principal value since numerator = 0
when $v_z = v_j$.

d. Penrose Condition applies for a distribution function with any number of humps.

9. Graphical Interpretation of Penrose Condition



b. Integral is ~~summation~~ of distribution above $F_0(v_j)$ (XXXX) minus that below $F_0(v_j)$ (VVV)

Weighted by function $\frac{1}{(v_z - v_j)^2}$



c. Thus, humps above minimum of $F(v_j)$ must be large enough that integral is positive.

Lecture #10 (Continued)

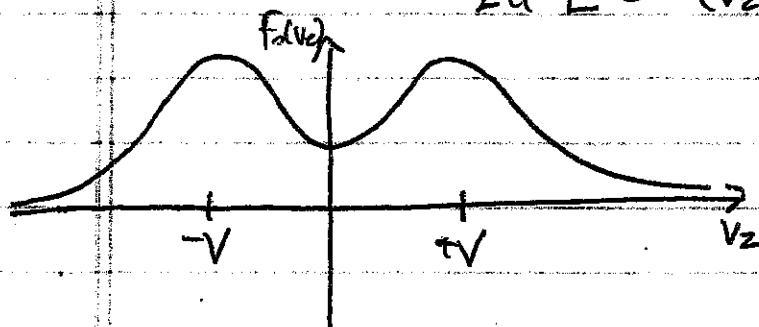
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I. (Continued) Example:

E. Counter-Streaming Beam Instability

1. DEF: Counter-Streaming Cauchy Distribution

$$F_0(v_z) = \frac{C}{2\pi} \left[\frac{1}{C^2 + (v_z - V)^2} + \frac{1}{C^2 + (v_z + V)^2} \right]$$



a. For $\lim_{C \rightarrow 0} F_0(v_z)$ we get two counter-streaming delta function beams. The "zero" temperature limit. UNSTABLE

- b. As C increases, eventually the distribution transitions to a single hump at $C = \sqrt{3}V$ STABLE by Gardner's Thm.
- c. Thus, at some point between $C=0$ and $C=\sqrt{3}V$ the system goes from unstable to stable.
(For increasing temperature, the distribution becomes stable).

2. Apply Penrose Condition

a. At peaks at $v_z = \pm V$, Penrose Condition is clearly negative.

b. At $v_z = 0$, we can show that

$$\int_{-\infty}^{\infty} dv_z \frac{F_0(v_z) - f_0(0)}{(v_z - 0)^2} = \frac{V^2 - C^2}{(V^2 + C^2)^2}$$

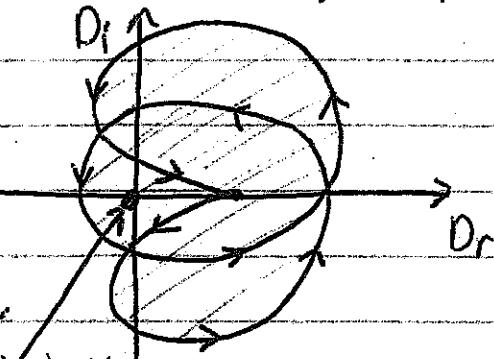
c. Thus Plasma is unstable when $V^2 - C^2 > 0$; or $V > C$

Unstable

3. Myguse Criterion

a. We can also evaluate D_r and D_i for this distribution

to show the $\gamma=0$ curve gives an unstable nose for $C < V$.



Unstable since $D=0$ is inside

II. Overview of Fluid vs. Kinetic Instabilities

A. Fluid vs. Kinetic Instabilities:

1. The two-stream instability of cold beams is a Fluid instability, because all particles react in the same way (no thermal spread of velocities).

2. Such fluid instabilities can be studied by fluid equations

3. Physical Picture of Two-Stream Instability

a. Consider a beam of positive charges streaming through a neutralizing background at rest.

b. Conservation of number density means $N; V_i = \text{constant}$.

c. If a ~~per~~ perturbation leads $V_i \uparrow$

to a decrease in beam velocity V_i at same point, the density

of ions must increase.

d. The resulting electric field

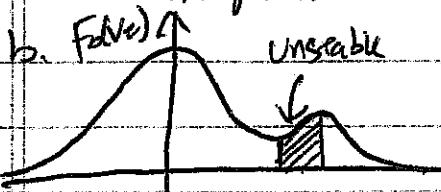
leads to a further slowing of

the beam \Rightarrow Positive feedback \Rightarrow UNSTABLE

e. Such an instability will continue to grow until some previously neglected nonlinear term halts the growth.

4. Kinetic Instability

a. The free energy in a finite temperature distribution function can lead to growth of instability due to interaction with resonance particles.



Only resonance particles are affected

c. Eventually free energy is tapped and kinetic instability

Settles