

## Lecture #2 - Single Particle Motion

### I. Overall Framework of Plasma Physics

Lorentz Force Law

$$m \frac{d\mathbf{x}_s}{dt} = q_s (\mathbf{E} + \mathbf{v}_s \times \mathbf{B})$$

Particles: Ions & Electrons

$$\left. \begin{array}{l} \text{position } \mathbf{x}_s \\ \text{velocity } \mathbf{v}_s \end{array} \right\} \Rightarrow \mathbf{r}_s$$

B & E Fields

Maxwell's

Equations

$$\nabla \cdot \mathbf{E} = \frac{\rho_e}{\epsilon_0}$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{j} + \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$

$$\nabla \cdot \mathbf{B} = 0$$

Single Particle Motion Description

What we want to study is how charged particles move in prescribed E & B fields.

### II. Larmor Motion: Constant, Uniform B with E=0

A. 1. Nonrelativistic Limit  $v \ll c$

2. Drop subscript "s" for species

3. Thus, for  $E=0$ ,  $m \frac{d\mathbf{v}}{dt} = q \mathbf{v} \times \mathbf{B}$

4. Take  $\mathbf{B} = B_0 \hat{z}$  were  $B_0 = \text{const}$

B. Solution:

$$1. \frac{dv_x}{dt} = \frac{q B_0}{m} v_y \quad (1)$$

$$2. \frac{dv_y}{dt} = -\frac{q B_0}{m} v_x \quad (2)$$

$$\frac{dv_z}{dt} = 0 \Rightarrow v_z = \text{constant}$$

## II. $B_0$ (Continued)

Hawes ②

2. Define:

$$\text{Cyclotron Frequency: } \Omega = \frac{qB_0}{m}$$

3. To Solve: a. Take  $\frac{d}{dt} \textcircled{1}$  and substitute  $\textcircled{3}$

$$\frac{d^2 V_x}{dt^2} = -\Omega^2 V_x$$

b. General Solution:  $V_x = A e^{-i\Omega t} + B e^{i\Omega t}$

c. Apply Initial Conditions & Solve for A & B

i. Take  $V_x = V_1$ ,  $V_y = 0$  at  $t = 0$ .  $\Rightarrow A = B = \frac{V_1}{2}$

ii. Let  $V_z = V_{11}$  at  $t = 0$  also.

d. Thus,  $V_x = V_1 \cos \Omega t$

$$V_y = -V_1 \sin \Omega t$$

$$V_z = V_{11}$$

$$x = \frac{V_1}{\Omega} \sin \Omega t + x_0$$

e. Solve for position:  $\frac{dx}{dt} = v \Rightarrow v = \frac{V_1}{\Omega} \cos \Omega t + v_0$

$$z = V_{11}t + z_0$$

f. Define: Larmor Radius  $r_L \equiv \frac{V_1}{\Omega} = \frac{mv_1}{qB_0}$

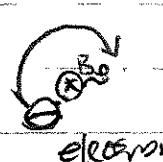
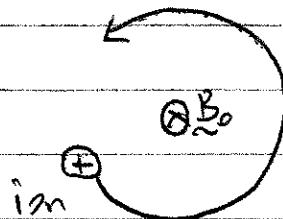
5. Summary:

a.  $\underline{x}(t) = r_L (\sin \Omega t \hat{x} + \cos \Omega t \hat{y}) + V_{11}t \hat{z} + \underline{x}_0$

b.  $\underline{v}(t) = V_1 (\cos \Omega t \hat{x} - \sin \Omega t \hat{y}) + V_{11} \hat{z}$

C. Properties

1. Diamagnetic:



Field due to Larmor motion opposes mean field

## II.C. (Continued)

Hawes ③

### 2. Conseant Energy:

$$a. \frac{dE}{dt} = \frac{d}{dt}\left(\frac{1}{2}mv^2\right) = \underline{v} \cdot \left(m \frac{d\underline{v}}{dt}\right) = \underline{v} \cdot [q(\underline{v} \times \underline{B})] = 0$$

b. Thus,  $\underline{v}_1 = \text{constant.}$

## III. $E \times B$ Drift: Conseant, Uniform $B$ and $E$

### A. Drift Motion $\underline{B} = B_0 \hat{z}$

$$1. m \frac{d\underline{v}}{dt} = q(E + \underline{v} \times \underline{B})$$

2. What velocity  $\underline{v}$  leads to  $\text{RHS} = 0?$   $\Rightarrow$  no acceleration  $\Rightarrow$  drift

$$a. E = -\underline{v} \times \underline{B}$$

$$b. \text{Cross with } \underline{B}: E \times \underline{B} = -(v \times B) \times B = B_0^2 (v - v_z \hat{z})$$

$$c. \text{Thus, } \underbrace{\underline{v} - v_z \hat{z}}_{\text{Perpendicular to } B_0} = \frac{E \times \underline{B}}{B_0^2}$$

### 3. Define

$$\boxed{\text{"}E \times B\text{" velocity } \underline{v}_E = \frac{E \times \underline{B}}{B_0^2}}$$

### B. Motion in $E \times B$ drift frame

$$1. \text{Solve for velocity } \underline{v} \text{ in } E \times B \text{ frame: } \underline{v} = \underline{v} + \underline{v}_E$$

$$2. \text{Substitute for } \underline{v}: m \frac{d\underline{v}}{dt} + m \frac{d\underline{v}_E}{dt} = q(E + \underline{v}_E \times \underline{B} + \underline{v} \times \underline{B})$$

$$a. \underline{v}_E \times \underline{B} = \frac{(E \times \hat{z}) \times \hat{z} B_0}{B_0^2} = E_z \hat{z} - E$$

$$b. \text{Thus } m \frac{d\underline{v}}{dt} = q(E_z \hat{z} + \underline{v} \times \underline{B})$$

$$3. \text{Parallel Motion } (\hat{z}): m \frac{dv_z}{dt} = qE_z \Rightarrow \boxed{v_z = \frac{qE_z}{m} t + v_{z0}}$$

### III B. (Continued)

Hanes ④

4. Perpendicular Motion:  $\underline{v}_\perp = v_\perp \hat{z}$

a.  $m \frac{d\underline{v}_\perp}{dt} = q(\underline{v}_\perp \times \underline{B})$  This is identical to the case with  $E=0$ .

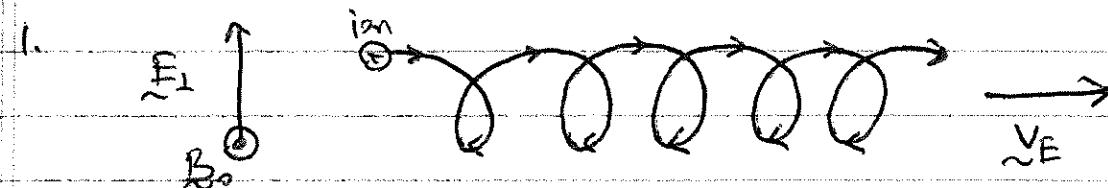
b. Thus,  $v_\perp = V_\perp (\cos \Omega t \hat{x} - \sin \Omega t \hat{y})$

c. In the  $\underline{E} \times \underline{B}$  drift frame, you have the usual Larmor motion

### 5. Full Solution:

$$\underline{v} = \underbrace{\left( \frac{q E_z t + v_{z0}}{m} \right) \hat{z}}_{\text{Parallel Motion}} + \underbrace{V_\perp (\cos \Omega t \hat{x} - \sin \Omega t \hat{y})}_{\text{Larmor Motion}} + \underbrace{\left( \frac{\underline{E} \times \underline{B}}{B_0^2} \right)}_{\text{E} \times \underline{B} \text{ drift}}$$

### C. Physical Picture:



- a. Acceleration by  $E \Rightarrow r_L$  increases } This asymmetry leads  
 Deceleration by  $E \Rightarrow r_L$  decreases } to the drift

2.  $\underline{E} \times \underline{B}$  drift is independent of charge.  $\Rightarrow$  No net current due to the  $\underline{E} \times \underline{B}$  drift.

### IV Multiple Timescale Methods

- A. 1. A powerful approach to solving many plasma physics problems is the use of multiple timescale methods.

2. In many problems, different components of the motion occur on disparate timescales.

## IV. A. 2. (Continued)

Hours (5)

a. For Example,  $E + \vec{B}$  drift

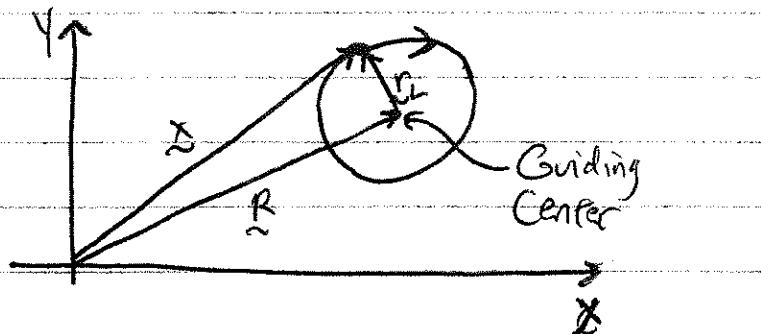
Decomposition  
of Motion:

- i. Rapid Larmor motion about field line
- ii. Slow drift across field line.

3. Define: Guiding Center

a. Position can be split into  
Guiding Center  $\underline{R}$   
plus Larmor motion  $\underline{r}_L$

$$\underline{x} = \underline{R} + \underline{r}_L$$



4. Basic concept for multiscale methods:

a. Average over fast timescale motion:

$$\int_0^{\frac{T}{2}} dt \underline{r}_L(t) = 0$$

b. This leaves the slow timescale drift motion  $R(t)$ .

## V. $\nabla B$ & Curvature Drifts: Constant, Non-uniform $B$ fields

A. In the fusion program, magnetic fields used to confine the plasma are neither straight nor uniform. We want to understand particle motion in  $B$  fields of varying strength and curved  $B$  fields.

B. Drift due to a General Force  $\underline{F}$

1. Analogous to  $E + \vec{B}$  drift, take  $m \frac{d\underline{v}}{dt} = q \underline{v} \times \underline{B} + \underline{f}$

$$\Rightarrow \underline{v}_D = \frac{1}{q} \frac{\underline{E} \times \underline{B}}{B_0^2}$$

2. NOTE: The direction of this drift depends on the charge sign.

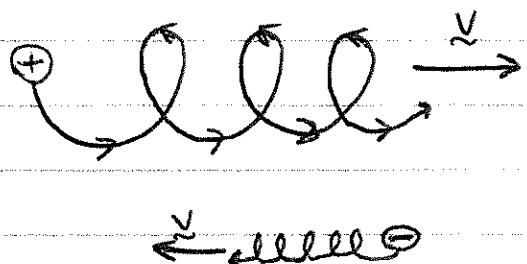
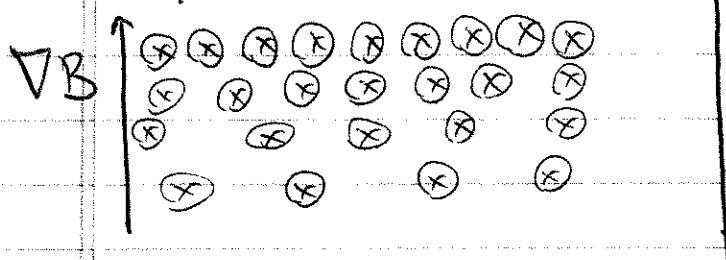
## II. (Continued)

Homework 6

### C. The $\nabla B$ ("GradB") Drift

1. Simple Case  $\nabla B \perp B$

2. Physical Picture:



a. Stronger  $B \Rightarrow$  smaller  $r_L$

Weaker  $B \Rightarrow$  larger  $r_L$

3. Multiscale Approach: a. Small scale: Larmor Radius  $r_L = \frac{v_1}{\omega}$

b. Large scale: B Scale length  $L \equiv (\frac{\nabla B}{B})^{-1}$

c. We may use a perturbative approach in the small expansion parameter  $\epsilon = \frac{r_L}{L} \ll 1$

d. We may derive the average force on the particle  $\langle F \rangle$   
(averaged over the Larmor period  $T = \frac{2\pi}{\omega}$ ) due to the  $\nabla B$ .

$$\langle F \rangle = -\frac{q}{2\Omega} \nabla B$$

4. The result is the  $\nabla B$  drift,

$$\nabla B = -\frac{v_1^2}{2\Omega} \frac{\nabla B \times B}{B^2}$$

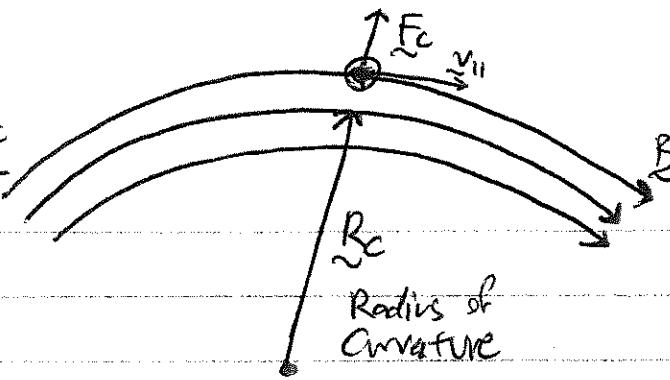
a. NOTE: Since  $\Omega = \frac{qB}{m}$ , the  $\nabla B$  drift depends on charge.  
⇒ Ions and electrons drift in opposite directions

b. Drift magnitude depends on perpendicular energy  $\frac{1}{2} m v_1^2$ .

## V. (Continued)

### D. Curvature Drift

#### 1. Physical Picture



Hawes ⑦

#### 2. Simple Estimate:

- a. For a particular moving along a circular path along  $\vec{B}$ , the centrifugal force felt by the particle is

$$\vec{F}_c = \frac{mv_{||}^2}{R_c} \hat{r} = \frac{mv_{||}^2}{R_c^2} R_c \hat{B}$$

- b. Treating this as the general force  $\vec{F}$ , we find

Curvature Drift

$$v_c \equiv \frac{mv_{||}^2}{qB^2} \frac{R_c \times \vec{B}}{R_c^2} = \frac{v_{||}^2}{\Omega B} \frac{R_c \times \vec{B}}{R_c^2}$$

- 3. Properties: a. Depends on parallel energy  $\frac{1}{2}mv_{||}^2$

- b. Again, ions & electrons drift in opposite directions

- 4. NOTE: When  $B$  field lines are curved, there is typically also a gradient in  $|B|$ , so both  $\nabla B$  & curvature drifts will be important.

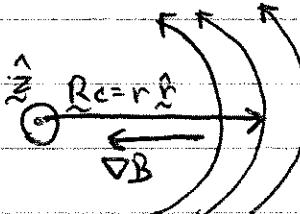
### E. Example: Current Carrying Wire

Consider a wire carrying a current  $I = I_0 \hat{z}$



- 1. In cylindrical coordinates  $(r, \phi, z)$ ,  $\vec{B} = \frac{\mu_0 I_0}{2\pi r} \hat{\phi}$

- 2. End on view



From NRL Plasma Formulary p.6,

$$\nabla B = \frac{\partial B}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial B}{\partial \phi} \hat{\phi} + \frac{\partial B}{\partial z} \hat{z} = \frac{-\mu_0 I_0}{2\pi r^2} \hat{r}$$

## V. E. (Continued)

Hawes (8)

### 3. DB Drift:

$$V_{DB} = -\frac{v_{\perp}^2}{2\Omega} \quad \frac{DB \times B}{B_0^2} = -\frac{v_{\perp}^2}{2\Omega} \quad \frac{\left(-\frac{\mu_0 I_0}{2\pi r^2} \hat{r}\right) \times \left(\frac{\mu_0 I_0}{2\pi r^2} \hat{z}\right)}{\left(\frac{\mu_0 I_0}{2\pi r}\right)^2} = +\frac{v_{\perp}^2}{2\Omega r} \hat{z}$$

### 4. Curvature Drift:

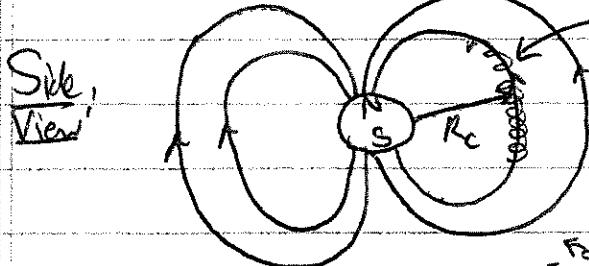
$$V_C = \frac{v_{\parallel}^2}{2B} \frac{B_0 \times B}{R_c^2} = \frac{v_{\parallel}^2}{2\left(\frac{\mu_0 I_0}{2\pi r}\right)} \frac{\left(r \hat{r}\right) \times \left(\frac{\mu_0 I_0}{2\pi r} \hat{z}\right)}{r^2} = \frac{v_{\parallel}^2}{\Omega r} \hat{z}$$

$$5. \text{ Net Drift: } V = V_{DB} + V_C = \frac{1}{\Omega r} \left( \frac{v_{\perp}^2}{2} + v_{\parallel}^2 \right) \hat{z}$$

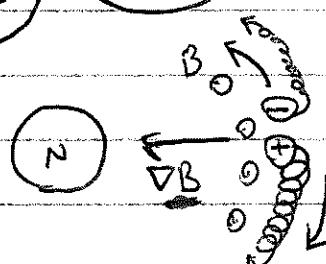
a. NOTE:  $\frac{1}{\Omega r} = \frac{m}{qB_r} = \frac{m 2\pi r}{q \mu_0 I_0 r}$ , so  $V = \frac{2\pi}{q \mu_0 I_0} \left( \frac{mv_{\perp}^2}{2} + mv_{\parallel}^2 \right) \hat{z}$

Velocity is independent of  $r$ !

## F. Example: Earth's Magnetosphere



Top View:



1. Particles trapped in Earth's dipole field experience DB & Langmuir drifts

2. These drifts produce the "ring current" in the westward direction

3. Strength of ring current is proportional to the energy of the particles.  
⇒ Magnetic Storms!