

- 1) Radiation dominated $< 47 \text{ kyr}$
 Matter dominated $47 \text{ kyr} - 9.8 \text{ Gyr}$
 Dark energy dominated $> 9.8 \text{ Gyr}$

- 2) Model universe as homogeneous sphere with mass M , radius r . Consider test mass m at surface.
 Energy is

$$E = K - U = \frac{1}{2}mv^2 - \frac{GMm}{r}$$

Critical case is dividing line between expanding/collapsing or unbound/bound when $E = 0$

$$E = 0 = \frac{1}{2}mv^2 - \frac{GMm}{r} \Rightarrow v^2 = \frac{2GM}{r} \quad (*)$$

mass of sphere $M = \rho \frac{4}{3} \pi r^3$ $\rho = \text{density}$

velocity of test mass $v = H_0 r$

where $H_0 = \text{Hubble constant}$

plug into (*) $H_0^2 r^2 = \frac{2G}{r} \frac{4}{3} \pi r^3 \rho$

$$\rightarrow H_0^2 = \frac{8\pi}{3} G \rho$$

$$\rho = \frac{3}{8\pi G} H_0^2 = \text{critical density}$$

3) definition of red shift $z = \frac{\lambda_{obs} - \lambda_e}{\lambda_e}$ (*)

λ_e = emitted wavelength

λ_{obs} = observed wavelength

photon wavelength stretches with universe

$$\text{so } \lambda_{obs} = \frac{\lambda_e}{a}$$

for example photon emitted when $a = 0.1$
is observed today when $a = 1$ with
wavelength $10 \times$ as long.

Plug into (*)

$$z = \frac{\lambda_e/a - \lambda_e}{\lambda_e} = \frac{1}{a} - 1 = z$$

4) $\nu_r \propto a^{-4} \Rightarrow \nu_r(a) = \nu_r(a=1) \cdot a^{-4} = \nu_{r,0} \cdot a^{-4}$

$$z = \frac{1}{a} - 1 \Rightarrow a = \frac{1}{z+1}$$

$$\nu_r(z=1100) = \nu_{r,0} \cdot (z+1)^{-4}$$

$$= 0.26 \text{ MeV m}^{-3} \cdot (1100+1)^{-4}$$

$$= 3.8 \times 10^{-11} \text{ MeV} \cdot \text{m}^{-3}$$

$$5) \quad \ddot{a} = H_0^2 \left[-\frac{\Omega_{r,50}}{a^3} - \frac{\Omega_{m,50}}{2a^2} + \Omega_\Lambda a \right] \quad (*)$$

cosmological constant dominates so
 $\Omega_\Lambda \gg \Omega_r, \Omega_m$

$$(*) \rightarrow \ddot{a} = H_0^2 \Omega_\Lambda a \quad (**)$$

try $a = e^{kt}$

$$\dot{a} = k e^{kt} \quad \ddot{a} = k^2 e^{kt}$$

plug into (**)

$$k^2 e^{kt} = H_0^2 \Omega_\Lambda e^{kt}$$

solution for all t if $k = H_0 \sqrt{\Omega_\Lambda}$

exponential growth since $a = e^{kt}$

time constant $\frac{1}{k} = H_0^{-1} \Omega_\Lambda^{-1/2} \approx 176 \text{ yr}$

universe expands by a factor $= e^1$

in a time $\frac{1}{k} = 176 \text{ yr}$