

# The Milky Way

- Shape
- Globular cluster system
- Galactic coordinates
- Pulsating stars
- Size of the Milky Way

Accretion disks around stellar-mass black holes emit most of their light as

- A) Radio waves
- B) Infrared light
- C) Visible light
- D) X-ray



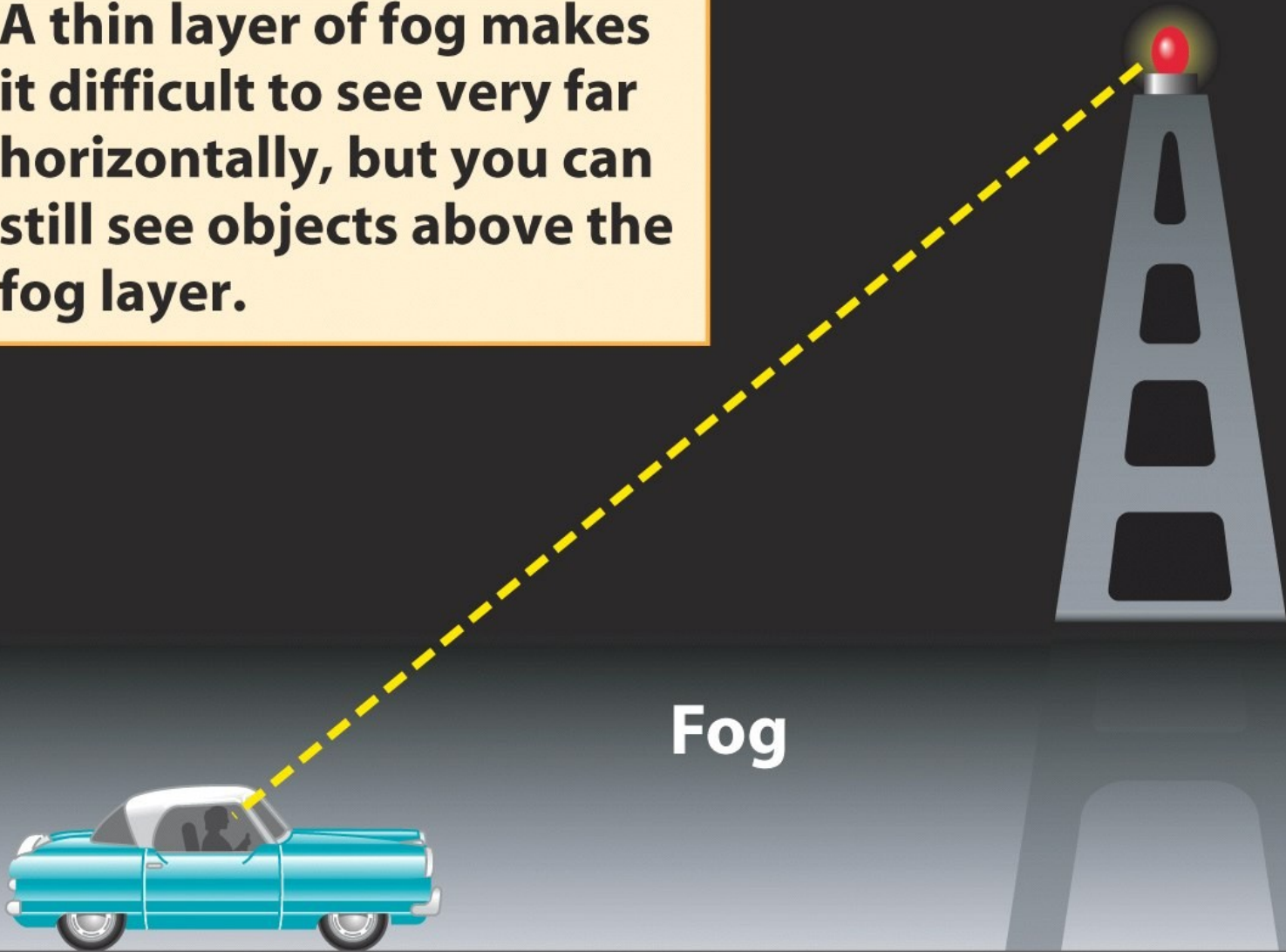




# How do we locate the center of the Milky Way?

- Can't see center directly with visible light because of obscuring clouds in the plane of the Galaxy
- Look above the plane of the galaxy

**A thin layer of fog makes it difficult to see very far horizontally, but you can still see objects above the fog layer.**



**Determining your position in the fog**

Interstellar dust makes it difficult to see very far within the plane of the Galaxy, but you can still see objects move above the galactic plane.

**Globular cluster**

**Sun**



**Earth**



**Disk of the  
Milky Way  
Galaxy**

**+ Center  
of Galaxy**

**Determining your position in the Galaxy**



M15

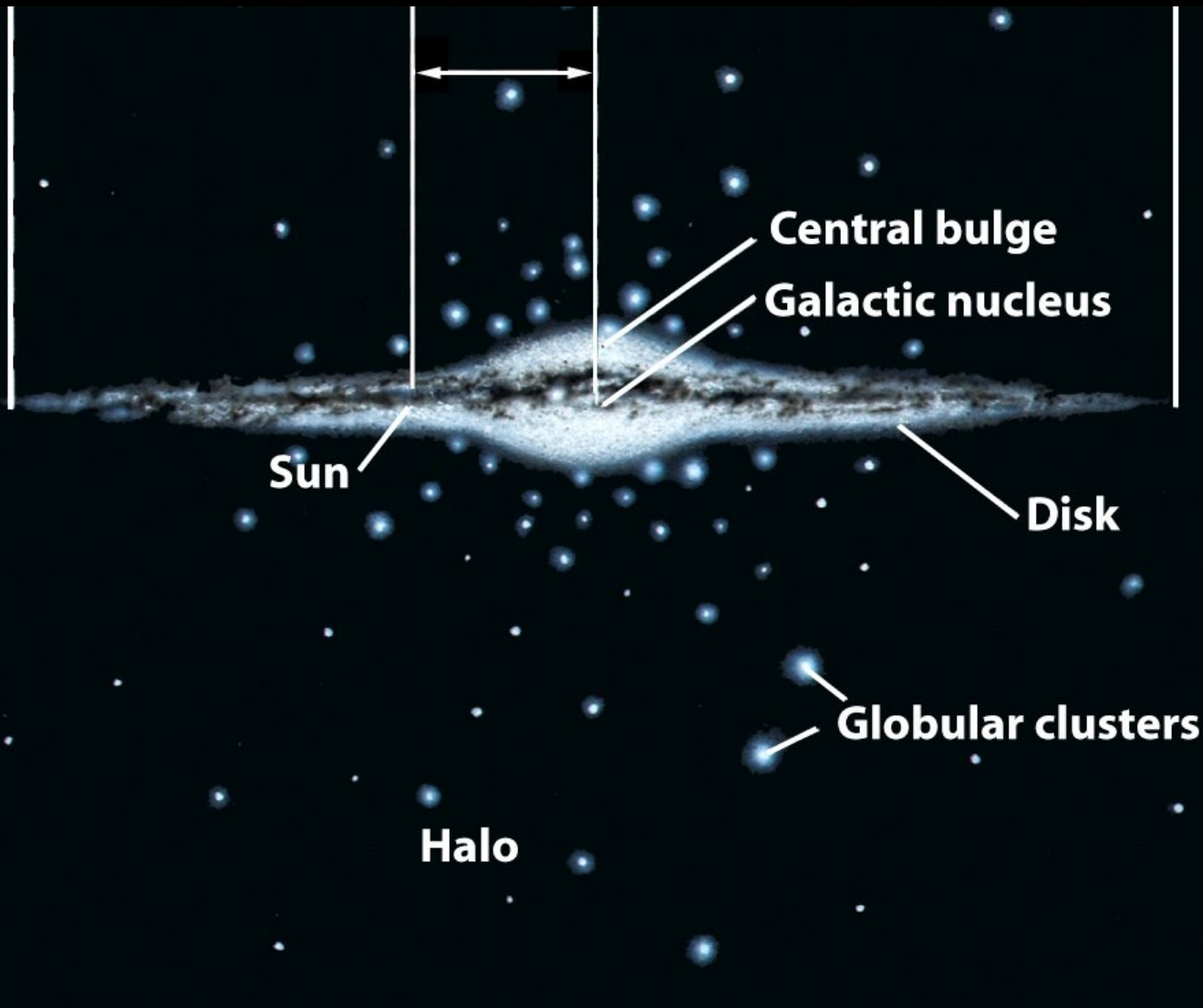


M13



# Globular clusters

- Compact, spherical group of stars
- Up to several 100,000 stars
- All stars formed together, same age
- Form a halo around the Milky Way



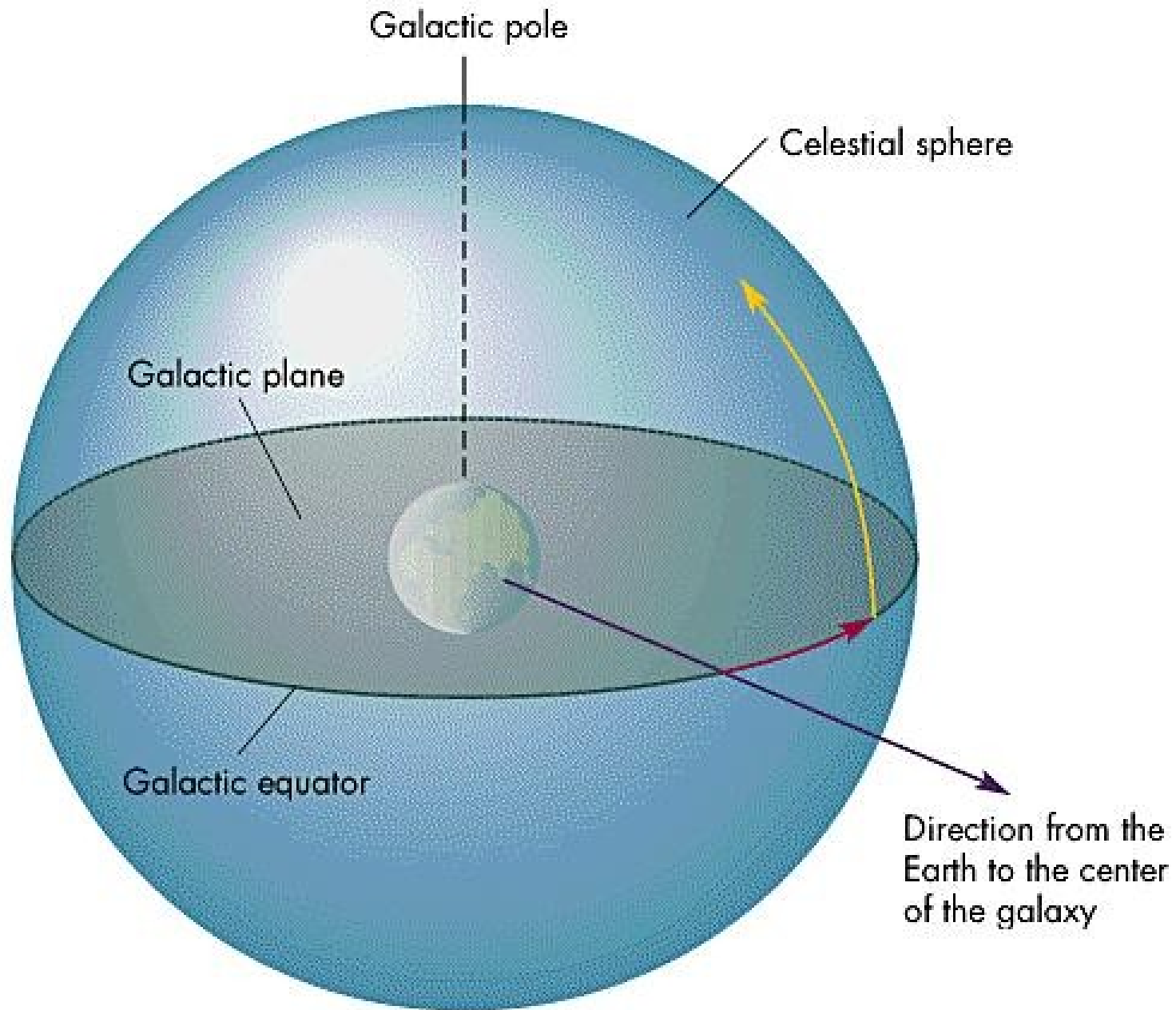
# Globular cluster system

- Centered on the center of the Milky Way
- Extends far above and below the plane
- By observing globular clusters, we can determine the direction to the center of the Milky Way and our distance from the center



What about globular clusters allows us to determine the Sun's distance from the center of the Galaxy?

- A) The clusters are confined to a thin plane.
- B) The clusters are all located at the same distance from the Sun.
- C) The center of the distribution of clusters is far from the Sun.
- D) All of the clusters are located in a region a few parsecs across.



Longitude ———  
Latitude ———



# But how do we find the distance to the Galactic center?

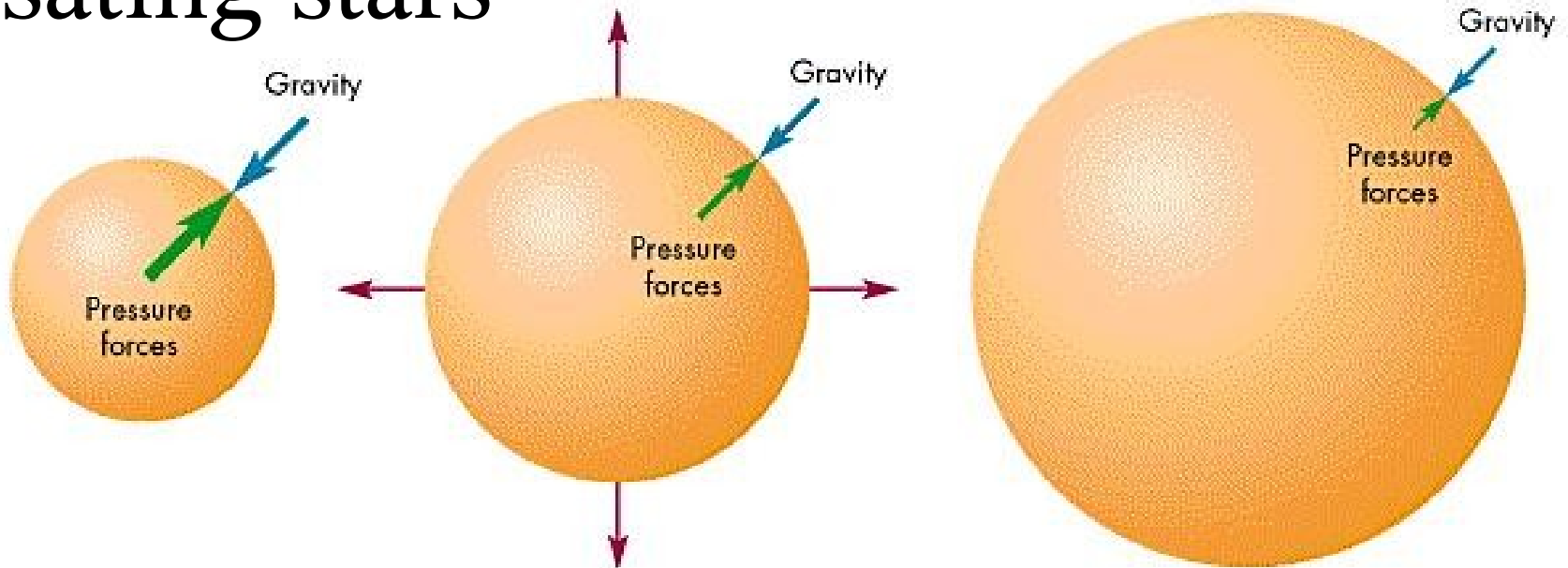
- We need the distances to the globular clusters
- Use pulsating stars in the clusters
- [Animation](#)

If a star is neither expanding nor contracting, we may assume that throughout the star there is a balance between pressure and

- A) temperature
- B) density
- C) luminosity
- D) gravity

Do mass on spring demo

# Pulsating stars

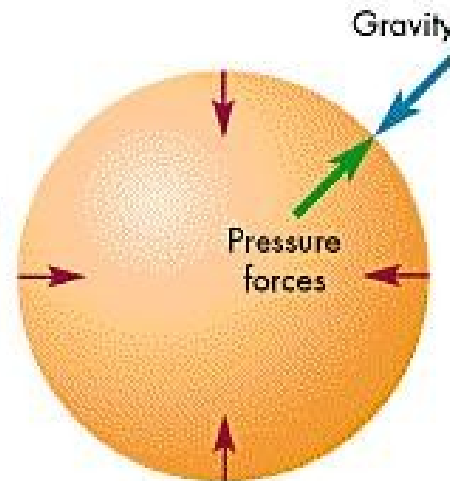


**A** Pressure forces exceed gravity: Pulsating star begins to expand

**B** Pressure and gravity balance but inertia makes the pulsating star expand further

**C** Gravity exceeds pressure: Pulsating star begins to contract

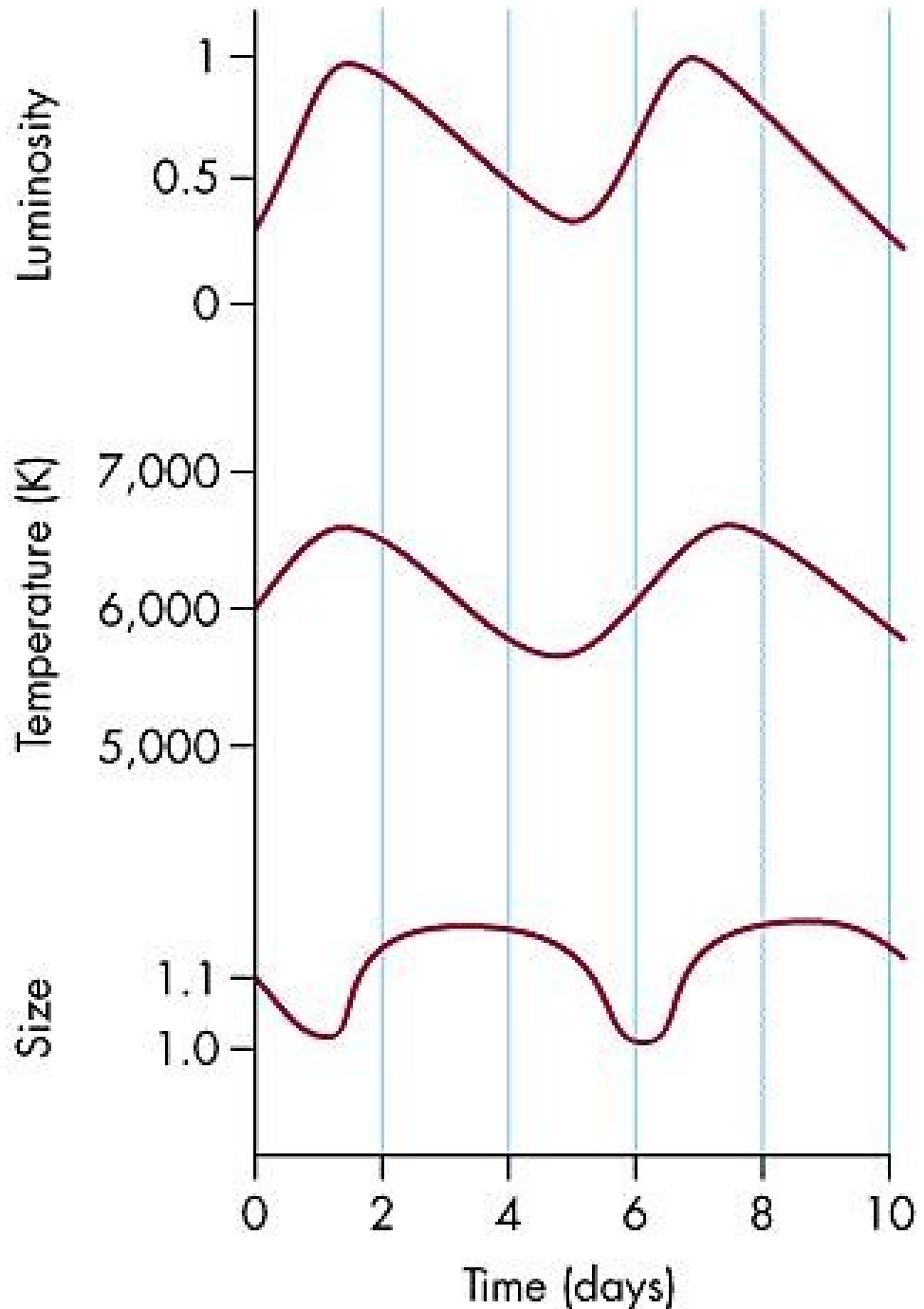
**D** Pressure and gravity balance but inertia makes the pulsating star contract further



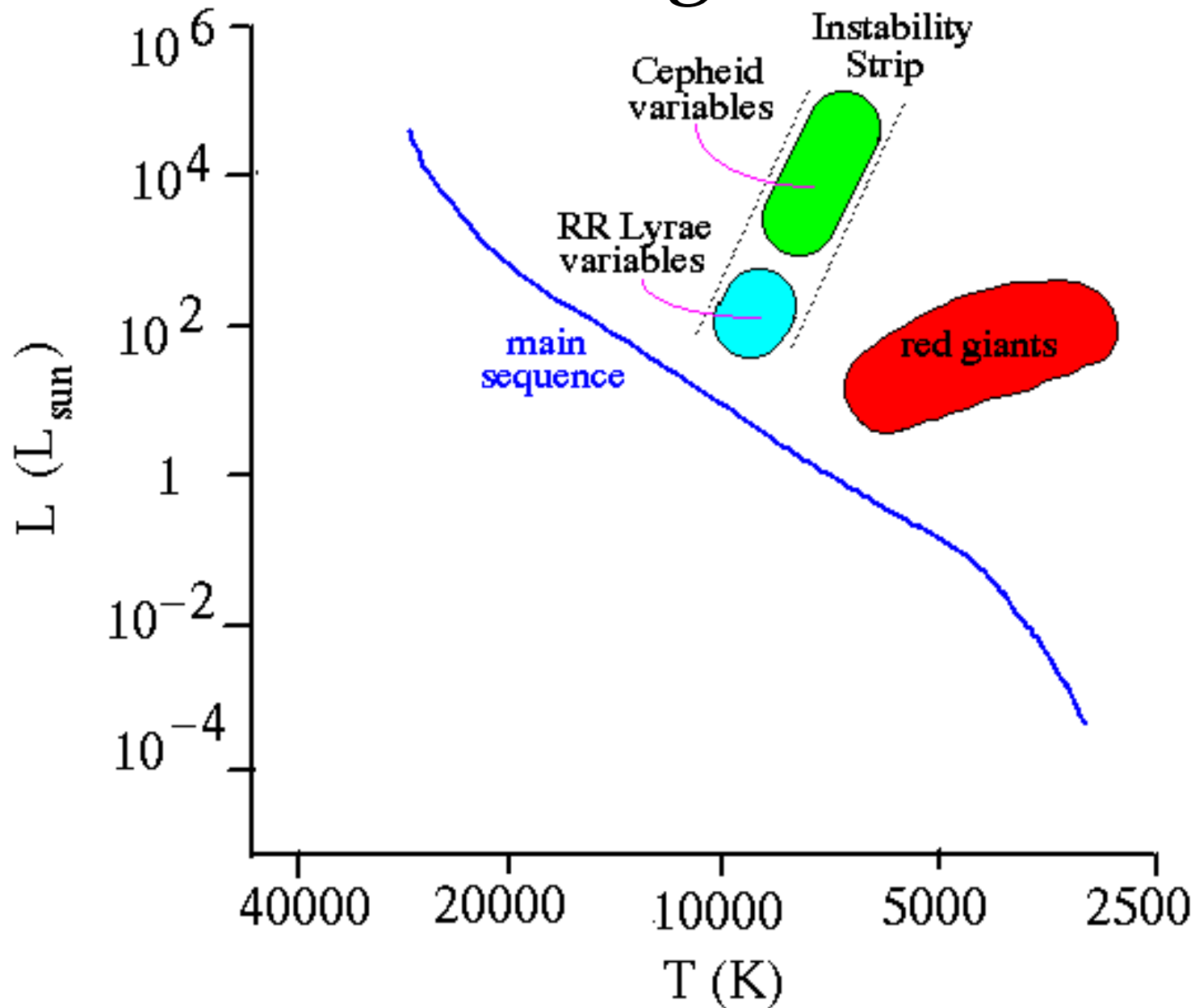
# Pulsation cycle

Rate of fusion in the core stays constant.

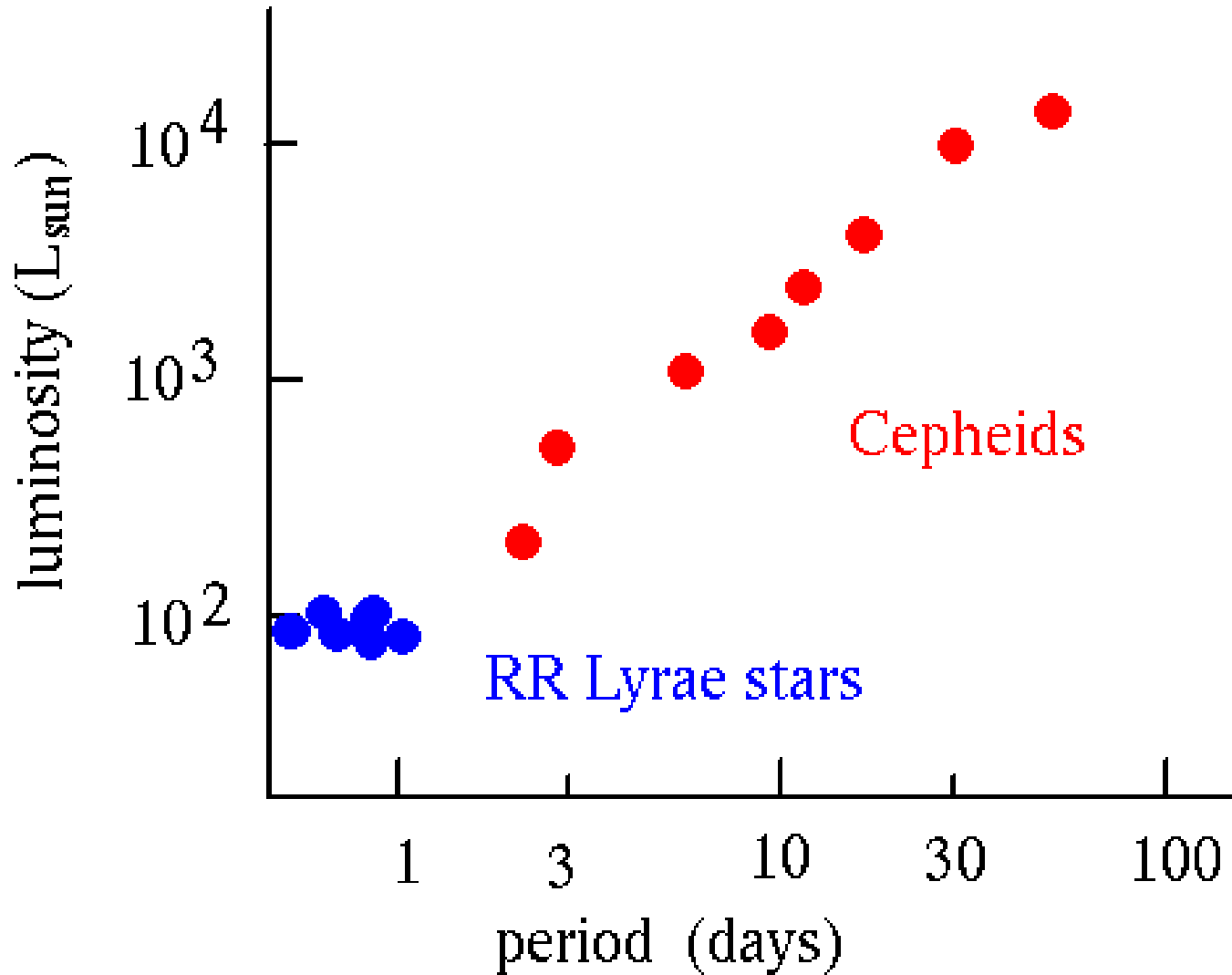
Transport of energy through outer layers of star oscillates.



# Pulsating stars



# Pulsating stars



What causes Cepheid variables to vary in brightness?

A) Eclipses

B) Changes in nuclear energy generation rate

C) Periodic changes in radius

D) Obscuration by clouds of dust

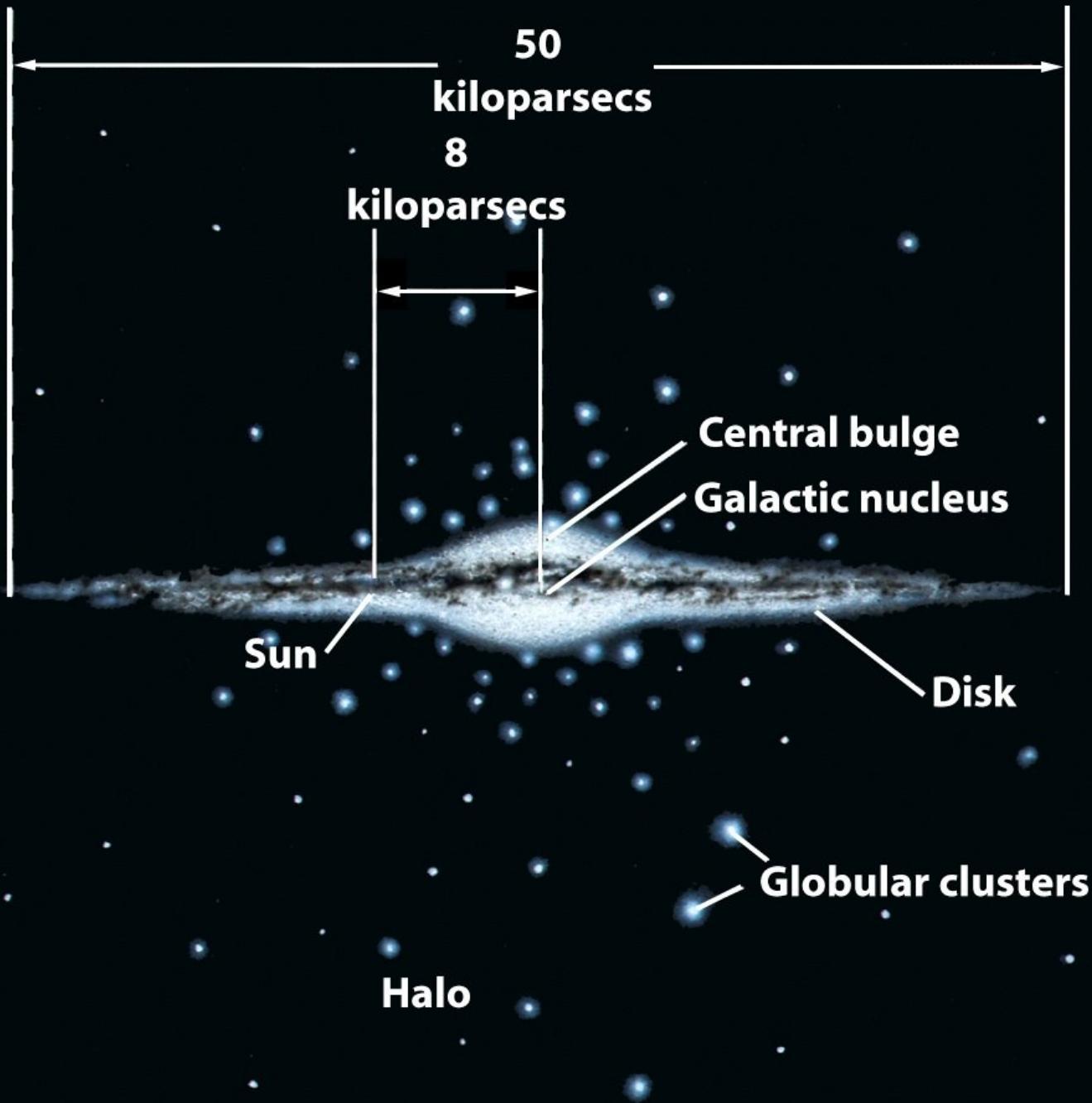
# Why is this useful?

Flux versus luminosity relation

$$\frac{\text{Flux}_A}{\text{Flux}_B} = \frac{\text{Luminosity}_A}{\text{Luminosity}_B} \left( \frac{\text{Distance}_B}{\text{Distance}_A} \right)^2$$

We can figure out the luminosity of a pulsating star by timing the pulsations. Since, we can measure its flux, we can then find the distance to the star.





# Review Questions

- What are globular clusters? How are they distributed in the Galaxy?
- What are Galactic coordinates?
- Why do some stars pulsate? Why are pulsating stars useful in measuring distances?
- What is the size and shape of the Milky Way?