



- The solution to F=ma for a planet is an ellipse with the Sun at one focus (Kepler's 1st Law)
- The semimajor axis and orbital period are related by:
   Kepler's 3rd Law (or is

$$P^{2} = \frac{4\pi^{2}a^{3}}{G(M+m)} \xrightarrow{\text{Keplet's 3rd}}_{\text{it?} ??????}$$

The application of Newtonian physics to orbital motion (continued)

Since the force is always in the direction of the center of the ellipse, the **torque** is always zero, and **angular momentum** is constant

Kepler's 2nd Law is a consequence

demonstration

Summary---Newton's laws of motion, and Newton's equation for the gravitational force (Newtonian mechanics) allow us to understand, and calculate with tremendous precision, the orbits of planets and other objects in the solar system.



- It is a prominent astronomical object
- Even a small telescope shows an incredible amount of detail on the surface
- It is the only astronomical object (other than the Earth) on which people have set foot.
- What we learned about the Moon helps us understand the whole solar system better





A consequence of synchronous rotation is that from Earth, we only see one side of the Moon



First Apollo landing here

The synchronous rotation of the Moon is a consequence of tides produced by the Earth (see textbook)







compared to the Earth

Masses (the other way of expressing how "big" an astronomical object is) and densities

- Moon, mass = 7.35E+22 kg, density = 3.34 g/cc
- Earth, mass = 5.97E+24 kg, density = 5.52 g/cc
- So the Moon has only 1.2 % of the mass of the Earth!
- To **really** keep things in perspective, consider the Sun, mass = 1.99E+30 kg

## Eclipses...we have already discussed, look over material on pp 184-191

Read over discussion of tides, the physical nature of them and their effects in the solar system

## Features of the lunar surface



With just your eye you can see the Maria and Terrae



Crater Tycho...lunar orbiter







Craters are the most distinctive features on the surface of the Moon. At one time they were thought to be unique to the Moon. Now we know them to be widespread and important solar system phenomena. Here are some important

- aspects. 1. They are large. There are 5 of them with diameters greater than 200
  - They are large. There are 5 of them with diameters greater than 200 kilometers.
     The diameters.
     The diameters of some of the famous ones are Tycho, 102 km, Polemaeus, 164km, Alphonsus, 108 km, Copernicus, 107 km, and Kepler, 31 km. We will see all of these during observing sessions of the Moon.
     By contrast, the famous Barringer Crater (or "Meteor Crater") in Arizona, which we formed its a come wire, here a diameters of 11 km.

  - which was formed in the same way, has a diameter of 1.1 km. It is now known that the lunar craters are *impact craters*, they were formed by the impact of massive meteors with the Moon. In the next lecture, we'll 4.
  - by the impact of master means with the most in the next rotate, we learn about when this happened.
    5. Many craters have central peaks, or mountains in the middle.
    6. A few, such as Tycho, Copernicus, and Kepler, have *rays*, or streaks of bright material pointing away from them.

The nature of the lunar craters was considered uncertain until rather late in the history of science. It was not until the early 1960s that it was considered established that they were impact craters, holes in the ground made by the explosion of a large object (rock) hitting the surface of the Moon. The diameter of the object is about 1/3 - 1/5 the diameter of the crater

