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## Other Worlds in Space: modern astronomy from the perspective of Exoplanets

Steven R. Spangler
Department of Physics and Astronomy
University of Iowa

What we are doing in this class: a survey of (much of ) contemporary astronomy, but with emphasis on an exciting, recent wrinkle: exoplanets


A meditation on the study of astronomy:
We want to feel, not just observe,
what the heavens offer..." John Goss, Astronomical League

## The motivating idea for this class

- We live on a planet (Earth) in the solar system
- All the planets orbit (and are controlled by) the Sun
- The Sun is a star
- The night sky is filled with stars
- Are there planets around some (or all) of those stars?
- Are some of those planets like the Earth?
- In 1995, we suspected the answer to (last 2 points) was yes, but we had no definite examples


## A mental goal of the class: an Earthlike planet around another star



## A sneak preview of where we are headed: the situation with exoplanets as of Dec. 2017

Exoplanet Discoveries

Total
confirmed
exoplanets
$=3,567$

Total
Kepler
$=2,525$


## Let's get back to basics: if we are going to

 understand exoplanets, we need to start with our own solar system

## "I seem to remember from junior high that there were others..."



## Jupiter, Saturn, Uranus, and Neptune are all further out

## The Solar System: in school we learn about the 8

 (or 9) planets that orbit the Sun


In solar system, the planets are where life is (at least in one case)

The Solar System is our best chance to study planets in detail (and learn some interesting astronomy, too)


The Earth as a Planet

## Main aspects of the Earth

- Surface temperature: about what is outside
- The atmosphere
- Liquid water



## Chemical composition of the Earth's atmosphere

- Nitrogen (N2): 78\%
- Oxygen (O2): 21\%
- Argon: 1\%
- Water vapor: 0-4\% !
- Carbon dioxide (CO2): 0.040 \%

By solar system standards, the Earth's atmospheric composition is unusual; the oxygen content is extremely unusual

A measure of how "thick" an atmosphere is: atmospheric pressure

- Atmospheric pressure is weight of overlying layers of air
- In physics has units of Force/(unit area)= Newtons/(square meter) = Pascals
- Atmospheric pressure at sea level = 1.013E+05 Pa
- Also called "one atmosphere"


## Liquid water on the surface of the Earth

The temperature and pressure at the surface of the Earth are close to the "Triple Point" of water (H2O), conditions where water simultaneously is in equilibrium in the solid, liquid and gaseous states.

This permits the hydrological cycle on Earth
The hydrological cycle and the triple point of water are almost certainly crucial for the existence of life on Earth

Amazingly, the Earth is not unique among solar system objects in having a "hydrological" cycle. Stay tuned.

## Overall Physical Structure of the Earth (the Earth as a ball)

- Mass = 5.98E+024
kilograms
- Diameter $=12756$ kilometers (7920 miles)
- Average density = 5.52 g/cc
- Denser than almost all rocks: heavy stuff in
 Earth's interior


## How do the other planets "stack up" next to the Earth?



## In our Solar System, the planets are so different, we have to put them in two groups

## Rough sizes of the terrestrial planets



This picture was known 50 years ago

## Where are Mercury and Venus in the night sky?



Very easy to find!

## Venus...closer and bigger. Expect more to see

## Venus 2004 Evening Apparition



## March 6

Phase =62\% $\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=19^{\prime \prime} .0$
Elong $=45^{\circ}$
App. Mag. $=-4.2$
$\Delta=0.88 \mathrm{AU}$


## March 29

Phase $=51 \%$
$\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=23^{\prime \prime} .6$
Elong $=46^{\circ}$
App. Mag. $=-4.4$
$\Delta=0.71 \mathrm{AU}$


April 10

## Phase $=44 \%$

 $\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=27^{\prime \prime} .2$ Elong $=45^{\circ}$ App. Mag. $=-4.4$$\Delta=0.62 \mathrm{AU}$


April 23
Phase =35\%

## Apparent $=$

 Diameter $=32^{\prime \prime} .6$Elong $=43^{\circ}$ App. Mag. $=-4.5$ $\Delta=0.52 \mathrm{AU}$


May 6
Phase $=24 \%$ $\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=39^{\prime \prime} .8$ Elong $=38^{\circ}$ App. Mag. $=-4.5$
$\Delta=0.42 \mathrm{~A} U$


May 15
Phase = 15\% $\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=46^{\prime \prime} .0$ Elong $=31^{\circ}$
App. Mag. $=-4.4$
$\Delta=0.36 \mathrm{AU}$


May 23
Phase =9\%
$\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=51^{\prime \prime} .8$
Elong $=\mathbf{2 3}{ }^{\circ}$
App. Mag. $=-4.3$
$\Delta=0.32 \mathrm{AU}$

May 29
Phase = 4\%
$\begin{aligned} & \text { Apparent } \\ & \text { Diameter }\end{aligned}=55^{\prime \prime} .4$
Diameter
Elong $=15^{\circ}$
App. Mag. $=-4.1$
$\Delta=0.30 \mathrm{AU}$

Current distance 1.61 AU, angular diameter 10.3 arcseconds. Why can't we see anything on the surface?

## Venus...Earth's twin



Venus is perpetually overcast. You cannot see the surface from Earth. Only the top of a cloud layer

## Similarities between Earth and Venus?

Fact that Earth and Venus are so similar in physical properties (size and mass), and that one cannot see the surface was conducive to imagination


## Seeing through the clouds of Venus with Radio Astronomy!



Radio astronomy and application of laws of physics allow us to measure the surface temperature of Venus (done first in the late 1950s)

- Surface temperature is 730 K
- That corresponds to 855 degrees Fahrenheit
- No longer "Earth's twin"
- What is responsible for this sort of temperature?
- Physics tells us!


## What the surface of Venus looks like



And now, a similar view of Venus, thanks to the Magellan orbiter


## Remaining questions about Venus

- Why is the surface temperature so high?
- Hint: it has a dense, carbon dioxide atmosphere
- Did Venus ever have oceans like the Earth?
- If it did have them, where did they go? (See article from Sky and Telescope, April 2008)


## Lessons of Venus for exoplanet studies

- A planet can be the "right size", and nearly in the "right place", but still be completely unsuitable for life


## Moving out in the solar system...Mars

View of Solar System from above
2018 MAR 20 15:30:00 UTC.
$2.0^{\circ}$ field of view

## Mars...start with the orbital parameters

- Semimajor axis: 1.524 AU
- Orbital Period: 1.88 years
- Eccentricity: 0.093
- Best opposition since 2003
 this summer (late July); take advantage of the chance to see Mars through a telescope


## Average physical characteristics of Mars



## What could we tell before the age of spacecraft and large telescopes?

Observations
showed features on surface ("canals of Mars")


## 1962 Mars Map

## Pre-space age telescopic observations yielded a number of intriguing features

- Seasonal variations in features
- Obliquity similar to Earth (25.0 degrees versus 23.5 degrees)
- Polar caps that change with the season
- Rotation period 24h 37m

- Clouds in an atmosphere

Led to speculation that Mars might be an Earthlike planet with primitive life (lichens)

By the 1960s it was known that the circa-1900 view of an Earth-like Mars with higher life forms could not be right; the atmosphere on the surface of Mars is about $0.7 \%$ the sea-level pressure on Earth. Water
 cannot exist in liquid form under those conditions.

The average surface temperature on Mars is also about 100 degrees
Fahrenheit below.

## The atmosphere of Mars: 95\% CO2

Most of the rest is nitrogen (diatomic molecular)

The chemical composition of Martian atmosphere is like Venus, not the current Earth
(But like the early Earth)

## Robotic Exploration of Mars



We have learned an incredible amount about Mars in the past few decades


An illustration of the fascinating results brought to us by robot spacecraft

A major result: apparent ancient water channels


Colorado River Canyon, Northern Arizona

Nanedi Vallis, Mars

## Determining the Ancient History of Mars

## $\frac{e^{2}}{2-20}$

## The central mystery of Mars since the 1970s

- The present climate of Mars is such that liquid water cannot last long
- Orbiting spacecraft show apparent evidence of flowing water channels (rivers and larger) in landscapes that formed 3 4 billion years ago
- How are these two (apparently) contradictory
 results to be reconciled?

These observations led to the idea of the "Warm, Wet Mars"

Mars 3.8 billion years ago?


Since 1990s, clear that the key is to land on the surface, do geology


The key to determining the ancient climate of Mars (and implications for life); robot geological
laboratories

## Curiosity



## The Mars Science Laboratory and Gale Crater

Clay deposits at bottom of central mountain


## The destination for "Curiosity"



## What have we learned from the MERs, "Mars Phoenix", and Mars Science Laboratory

- It now seems clear that there were standing bodies of water for long periods of time early in Martian history
- Evidence is presence of hematite, jarosite, and other minerals that form in lakes or oceans
- Relative absence of carbonate rocks is due to alternative chemistry in acidic water
- But, apparently oldest rock strata do contain carbonate rocks and clay (montmorillonite)
- Mars Science Laboratory has found abundant mineralogical evidence for water in Gale Crater
- Apparently there were liquid water bodies on Mars in at least some locations (Gale Crater) long enough for life to have developed.


## What happened to the atmosphere?

- The rocks found on the Martian surface required higher atmospheric pressures than presently exist
- Mars must have lost its atmosphere
- Can we find independent evidence for this? Yes!
Finding of the Maven spacecraft
- How long did it take to lose its atmosphere?---We'll have to wait for the answer

The El Dorado of future missions: Mars Sample Return

## Stay tuned



What about the more distant
planets in the Solar System?

What can they tell us about how planets formed, the nature of other solar systems?

## Exploring further out in the Solar System



Jupiter, Saturn, and friends

## Jupiter and Saturn: orbital characteristics

| Planet | $a(A U)$ | $P(y r s)$ | ecc | Incl (deg) |
| :--- | :--- | :--- | :--- | :--- |
| Jupiter | 5.20 | 11.9 | 0.049 | 1.3 |
| Saturn | 9.58 | 29.4 | 0.057 | 2.5 |



What do they look like? Let's start with Jupiter

## Jupiter...largest planet in the solar system

## Basic properties of Jupiter and Saturn

- Jupiter: 11.2 X diameter of Earth and 318 X mass
- Saturn: 9.5 X diameter of Earth and 95 X the mass
- Jupiter and Saturn: the "giant planets"
- Jupiter and Saturn are "all atmosphere", and mainly hydrogen and helium
*T. Jupiter and the Earth


## The future exploration of Jupiter... Juno (arrived

 last summer)

Launch: August 5, 2011... Arrival at Jupiter: July 4, 2016

## Juno Spacecraft



SANCECReFT DMENSONS
Dimeter 66 foxt 60 metiens
llught 15 fect ( 45 metion)

For more information missionjuno.swriedu \& wwwenasa.gow/uno


## Juno's Instruments

Gravity Sclence and Mugnetometers Study Jupter's drep structure by mapping the planet's gravity foid and magnetio field

Microwave Radiometer
Probe Jupiter's deep atrosphere and measure how much wastr (and hence oxygen) is there

JED JADE and Waves
Sample electric fields, plasma waves and particles around Jupiter to determine how the magratic feld is connected to the atmosphere.
and especially the auvoras (northern and scuthem lights)

UVS and JIRAM
Using utrawislet and infrared cameras, take imesas of the atrocsphere and ausoras. including chamical frgapents of the gases present

JuncCam
Take spectacular close-up, color imeges

## The Juno spacecraft is giving us new insights into Jupiter

There is really no surface on Jupiter; the gas density and pressure get larger and larger, matter weirder and weirder


## Are Jupiter and Saturn planets or stars?

Jupiter as
seen at
infrared
wavelengths


Jupiter emits $70 \%$ more radiation to space than it receives from the Sun. It has an "engine" inside

The Moons of Jupiter

## Jupiter has many moons

- 12 when I started studying astronomy
- A standard textbook lists 38
- Most important are the 4 Galilean satellites, lo, Europa, Ganymede, Callisto


## From Earth, it is difficult to learn too much about the Galilean satellites

At opposition of Jupiter, the angular diameter of Ganymede is 1.7 arcseconds

Pre space-age telescope observations revealed a little bit about size, reflectivity (albedo), and surface features

## Basic data on the Galilean satellites

| moon | a $(\mathrm{km})$ | $\mathrm{P}($ days $)$ | $\mathrm{D}(\mathrm{km})$ |
| :--- | :--- | :--- | :--- |
| Io | 422,000 | 1.769 | 3630 |
| Europa | 671,000 | 3.551 | 3130 |
| Ganymede | $1,071,000$ | 7.155 | 5280 |
| Callisto | $1,884,000$ | 16.689 | 4840 |
| MOON | 384,000 | 27.32 | 3476 |

The largest Moon, Ganymede, has a diameter about 50 percent larger than our Moon

## What are the moons of Jupiter like? Would do they look like "up close"?



The 1968 movie "2001 A Space Odyssey" has them similar and looking like our Moon

## All of them together (sibling portrait)



Images scaled to give correct relative sizes

## Exploring Europa



What is under its ice-covered plains?

Europa is slightly smaller and less massive than our Moon. It is of interest because the entire moon is encased in ice. There are cracks and other features that hint at liquid water at some point below the surface,

## Views of the cracks from Galileo



Picture about 100 miles on a side

## A related phenomenon. The ice rafts of

Similar features seen in arctic ocean and are due to flows of ocean underneath


## Evidence for flows from beneath the surface of Europa



There is evidence (circumstantial) for liquid water under the surface, but how far down is it? What is below the water?

Spacecraft missions to Europa over the next few decades might tell us

## Speculations on interior structure of Europa



## A summary of what we know about Europa

- Slightly smaller in mass and diameter than the Moon
- Surface covered with water ice casing
- Evidence for surface "activity" from cracks and grooves, and ice rafts
- Small numbers of craters implies surface has reformed in last 100 million years
- Estimates that liquid layer, "sealed ocean" is between 10-50 kilometers below the surface, with possible rocky sea floor


## A future Europa Lander could tell us much about the possible subsurface ocean of Europa



## Further out...to the edge of the Solar System



## Uranus and Neptune... where are they? Let's look at a Table!

| Planet | $\mathrm{a}(\mathrm{AU})$ | $\mathrm{P}(\mathrm{yr})$ | ecc | Incl <br> (degrees) |
| :--- | :--- | :--- | :--- | :--- |
| Uranus | 19.20 | 83.7 | 0.046 | 0.8 |
| Neptune | 30.5 | 163.7 | 0.011 | 1.8 |

Current locations: Uranus: Pisces; Neptune: Aquarius.

What do they look like?

## Our best views (and scientific information) come from visits (flybys) of the Voyager 2 spacecraft



## Uranus as seen by Voyager 2



## Neptune as seen by Voyager 2

They look like "blue Jupiters"


## How do they match up to Jupiter and Saturn?



Smaller than Jupiter and Saturn; much bigger than the Earth

## Uranus (and Neptune) substantially larger than Earth



## Uranus and Neptune (just the facts, Ma’am)



| planet | D (rel to Earth) | M (rel to Earth) |
| :--- | :--- | :--- |
| Uranus | 4.01 | 14.5 |
| Neptune | 3.88 | 17.1 |

Uranus and Neptune are relevant in the exoplanet context

## How do we summarize all of this?



- The small rock planets are in close to the Sun
- The big, gaseous planets are way out, far from the Sun
- Mars may have had habitable conditions on its surface billions of years ago
- Moons of the outer planets (Europa, Titan) may be promising locations for life

