

Lecture #25 The Holy Grail of Plasma Physics: Controlled Thermonuclear Fusion

Hawes ①

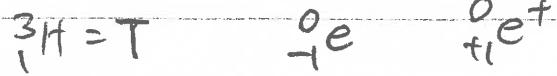
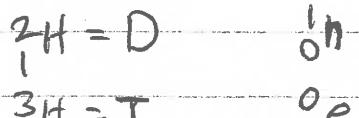
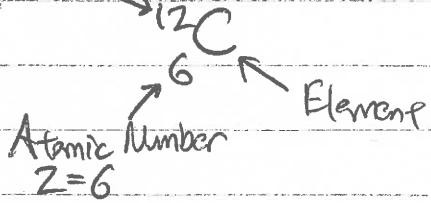
I. Nuclear Fusion: The Energy Source of the Stars

A. The CNO Cycle

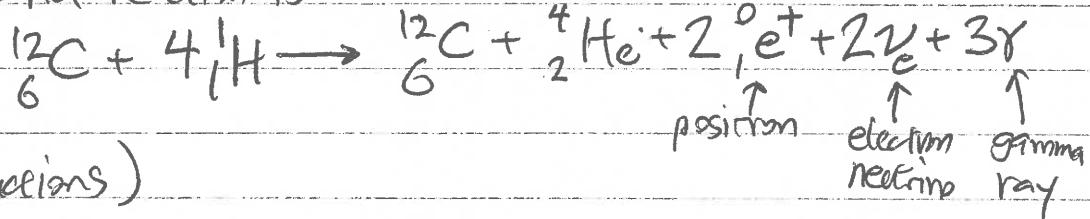
1. In 1938/1939, Carl von Weizsäcker & Hans Bethe independently proposed this chain of nuclear reactions as the source for stellar luminosity

2. $^{12}_{\text{C}}$ acts as a catalyst to fuse four protons into an alpha particle

a. Mass Number $A=12$



- b. Thus, the net reaction is



3. This is the dominant power source for stars with $M > 1.5 M_\odot$

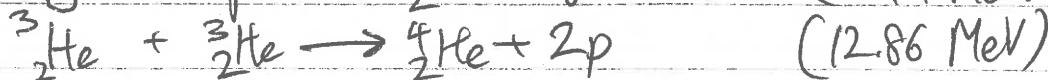
a. CNO reactions occur for $T > 1.3 \times 10^7 \text{ K}$

b. Our Sun has a central temperature $T \approx 1.57 \times 10^6 \text{ K}$

\Rightarrow only 1.7% of α 's are produced by CNO in Sun.

- c. At $T > 1.7 \times 10^7 \text{ K}$ CNO begins to dominate.

B. The p-p cycle



Lecture #25 (Continued)

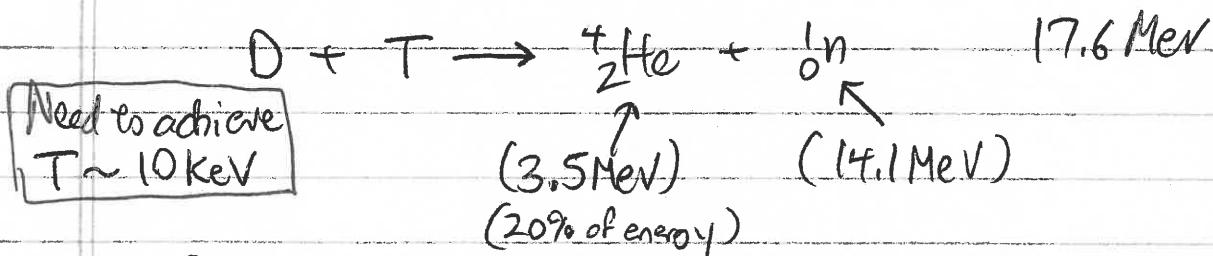
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T B. (Continued)

2. The lower central temperatures in the Sun lead to dominance of the pp cycle.
3. p-p cycle requires $T > 4 \times 10^6 \text{ K}$.

C. Laboratory Fusion: D-T reaction

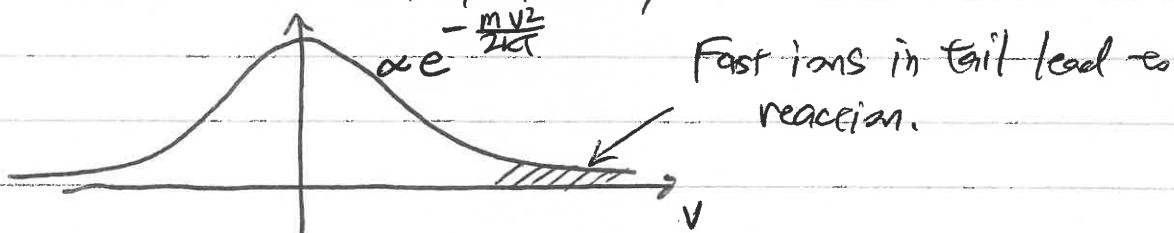
1. Basic characteristics for fusion in the lab:



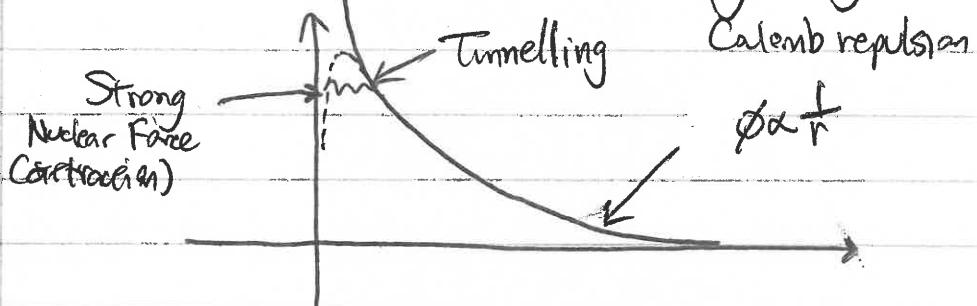
2. Reactions relevant to Laboratory Fusion \Rightarrow p. 44-45 NRL Plasma Formulary.

3. DT reaction cross-section peaks at $\sim 50 \text{ keV}$, but you do not need a plasma with $T \sim 50 \text{ keV} \sim 6 \times 10^8 \text{ K}$.

- a. For a Maxwellian equilibrium,



- b. Quantum Mechanical Tunnelling through Coulomb Barrier



4. Difficulties

- a. Confinement
- b. Fast neutrons

- c. Tritium Supply.

Lecture #25 (Continued)

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I. C. (Continued)

5. Neutrons: 14.1 MeV neutrons are not contained
 \Rightarrow Can cause severe damage and weakening of reactor material.

G. Tritium:

a. Cost for tritium is estimated at \$84,000 to \$130,000 / gram.

b. Need a way to produce tritium

\Rightarrow Breeder Reactions: Lithium Blanket



c. Can use neutrons from D-T reaction to produce more T.

d. World's supply of Lithium is limited, but sufficient to supply world's energy demands for thousands of years.

D. The Lawson Criterion:

1. Balance power lost by plasma with power released by fusion reaction.

2. Fusion power in α 's:

$$a. n_D = n_T = n$$

$$b. P_\alpha = \frac{1}{4} \underbrace{\langle \sigma v \rangle}_{\substack{\text{cross-section} \times \text{Velocity} \\ \text{averaged over Maxwellian distribution}}} n^2 E_\alpha \quad E_\alpha = 3.5 \text{ MeV}$$

$\langle \sigma v \rangle$
 averaged over Maxwellian distribution

c. Efficiency for retaining α power in plasma: η

3. Power Loss: a. Loss of confinement $P_C = -\frac{3nKT}{\tau}$

τ = confinement time

b. Radiative losses (for example, Bremsstrahlung from $P_B = \alpha n^2 T^{1/2}$ $\alpha = \text{const.}$ electron-ion collisions)

Lesson #25 (Continued)

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I.D. (Continued)

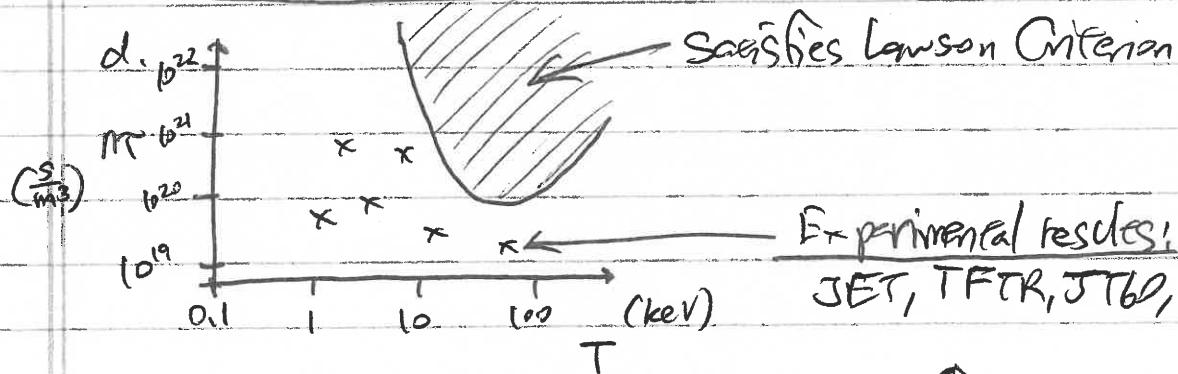
4. Power Balance: $P_x + P_B + P_c > 0$ to sustain fusion

$$a. \frac{q}{4} \langle \sigma v \rangle n^2 E_\alpha - \alpha n^2 T^{\frac{1}{2}} = \frac{3nKT}{T} > 0$$

$$b. \frac{q}{4} \langle \sigma v \rangle E_\alpha - \alpha T^{\frac{1}{2}} > \frac{3nKT}{n^2 \tau}$$

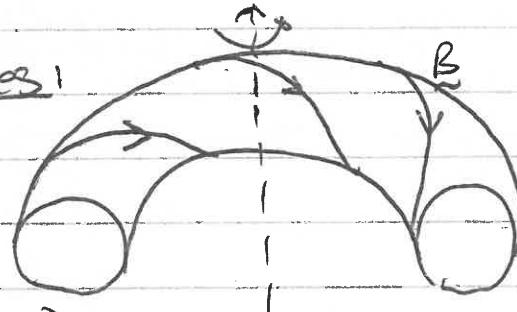
$$c. \boxed{n\tau > \frac{3KT}{\frac{q}{4} \langle \sigma v \rangle E_\alpha - \alpha T^{\frac{1}{2}}}}$$

Lawson Criterion



II. Plasma Confinement Schemes:

A. Magnetic Confinement:



1. Tokamak: Most successful.

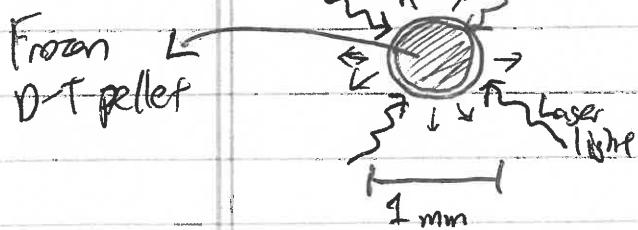
(Russian: Toroidal Magnetic Chamber)

2. Other: Z-pinch, Reverse Field Pinches, Stellarator, etc.

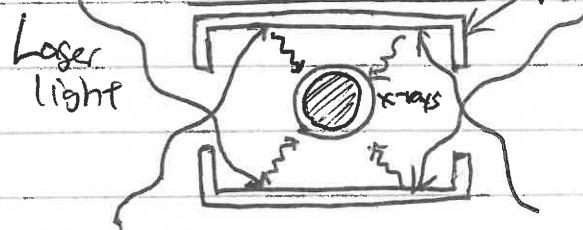
3. Requires plasma confinement times of at least a few seconds.

B. Inertial Confinement (Laser Fusion)

1. Direct Drive:



2. Indirect Drive: Hohlraum



3. Bomb reaction compresses plasma to 10^3 times liquid density for very short times.