

Physics II: 1702

Gravity, Electricity, & Magnetism

Professor Jasper Halekas

Van Allen 70 [Clicker Channel #18]

MWF 11:30-12:30 Lecture, Th 12:30-1:30 Discussion

Announcements I

- First homework (math and conceptual sections) due tonight at 11 pm on Wiley Plus

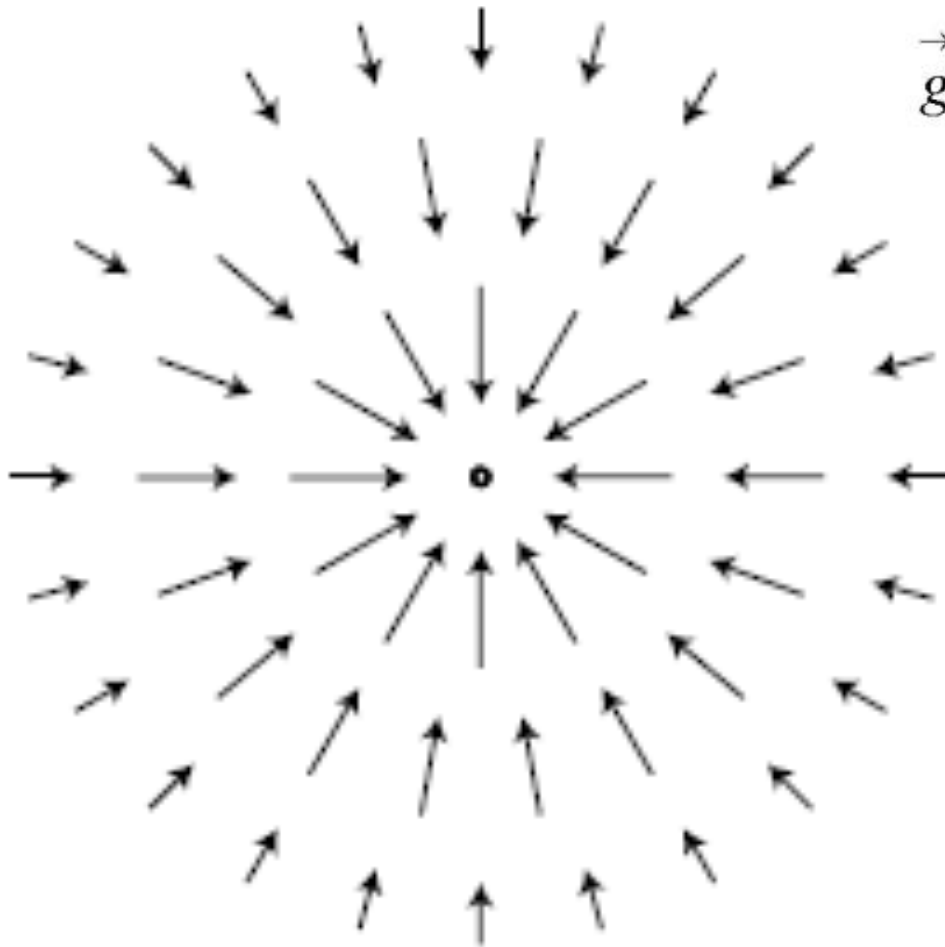
Announcements II

- University Hawk Shop temporarily out of lab manuals
- Iowa book may have some old copies (23 left as of Wednesday afternoon)
- Wherever you check, check under 1512, 1612, and 1702 (same lab manual for all)
- If none of these avenues pan out by Monday, I have a copy of the pre-lab questions and worksheets for E1 that I can send to you

Announcements III

- Please contact me if you are in the late lab section and wish to attend the caucuses on Monday night (which I highly encourage you to do, despite the fact that it will complicate our lives slightly!)

Gravitational Field

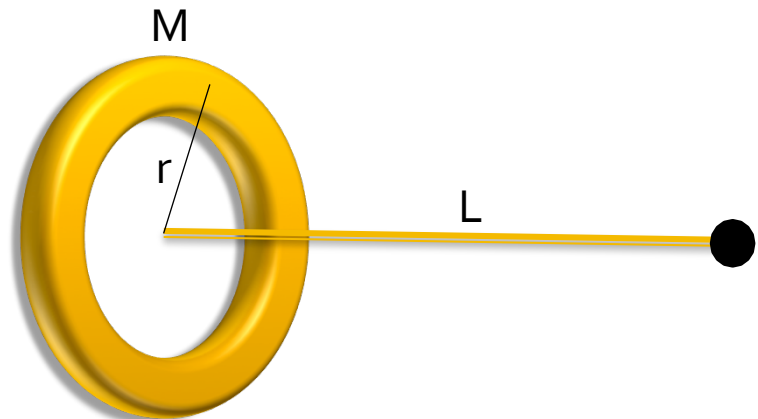


$$\vec{g} = \frac{\vec{F}_g}{m} = -G \frac{mM}{mr^2} \hat{r} = -\frac{GM}{r^2} \hat{r}$$

Superposition and Ring of Mass

- What is the gravitational field due to a thin ring of matter with radius r and total mass M , at a point a distance L along the line from the center of the ring?

- $-GML/(r^2 + L^2)^{3/2}$



- $1/(r^2 + L^2)$ from inverse square law
- Extra factor of $L/\sqrt{(r^2 + L^2)}$ because only axial components add (radial components cancel)

The principle of symmetry

- Since the ring is rotationally symmetric, nothing would change if we rotate the ring around its center
- This actually directly implies that the field at a point along the line from the center has to be in the axial direction
- To see this, imagine that it wasn't
 - This would imply that the field would rotate when you rotated the ring, but this clearly can't be since the ring is completely symmetric

Gravity of a spherical shell

- The gravitational field of a spherical shell can be calculated by just adding up the field of many little rings.
 - This turns out to be a rather painful integral (which I will spare you).
- The conclusion is somewhat remarkable
 - The field outside a shell of matter is equal to the field if the entire mass of the shell was at the center of the shell
 - The field inside a shell of matter is zero

Gravity at Earth's Surface

- By the shell theorem, the gravity at the Earth's surface is equivalent to the gravity from a point mass at the center of the Earth
- $M_E = 5.972 \times 10^{24} \text{ kg}$
- $r_E = 6378000 \text{ m}$ (at equator)
- $g = GM_E/r_E^2 = 9.792 \text{ m/s}^2$

Gravity Below Earth's Surface

- Knowing that the field outside of a shell of matter is proportional to its mass divided by the square of the distance from its center, and that the field inside a shell of matter is zero, can you predict how the field inside the Earth varies with radius r ?
 - A. g proportional to $1/r^2$
 - B. g proportional to $1/r$
 - C. g constant
 - D. g proportional r
 - E. g proportional to r^2

\vec{g} outside shell of mass M and radius r

$$= -\frac{GM}{r^2} \hat{r}$$

\vec{g} of a sphere is just that of many shells

$$\text{so } \vec{g} = -\frac{GM_{\text{inside}}}{r^2} \hat{r}$$

where M_{inside} is all the mass inside the radius r

What about inside a sphere?

All that counts is mass inside

$$M_{\text{inside}} = \rho V_{\text{inside}}$$

$$= \frac{M_{\text{tot}}}{V_{\text{tot}}} V_{\text{inside}}$$

$$= M_{\text{tot}} \frac{\frac{4}{3}\pi r^3}{\frac{4}{3}\pi R^3} = M_{\text{tot}} \frac{r^3}{R^3}$$

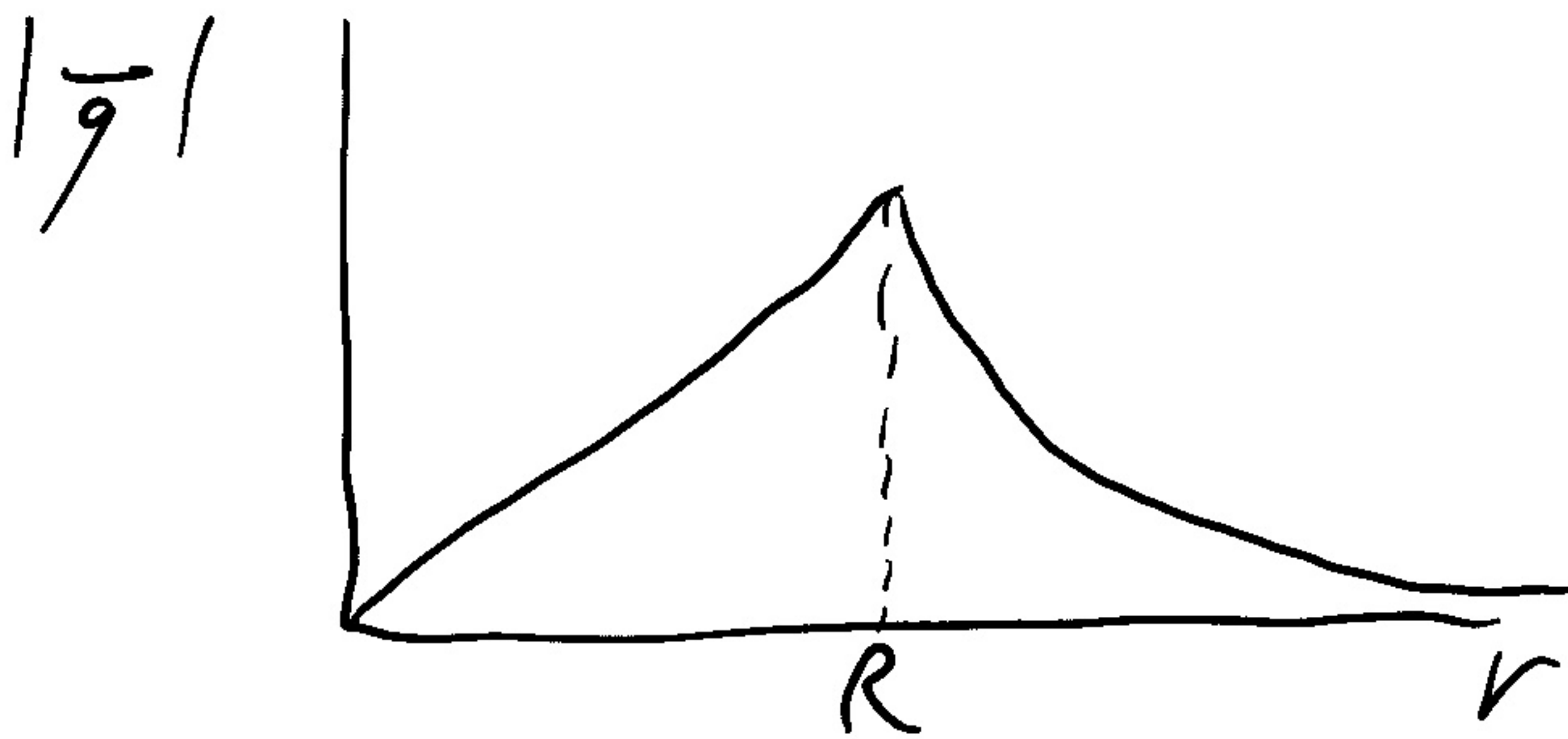
$$\vec{g} = -\frac{GM r^3/R^3}{r^2} \hat{r} = -\frac{GM r}{R^3} \hat{r}$$

Full solution:

$$\vec{g} = -\frac{GM}{R^3} \vec{r} \quad r < R$$

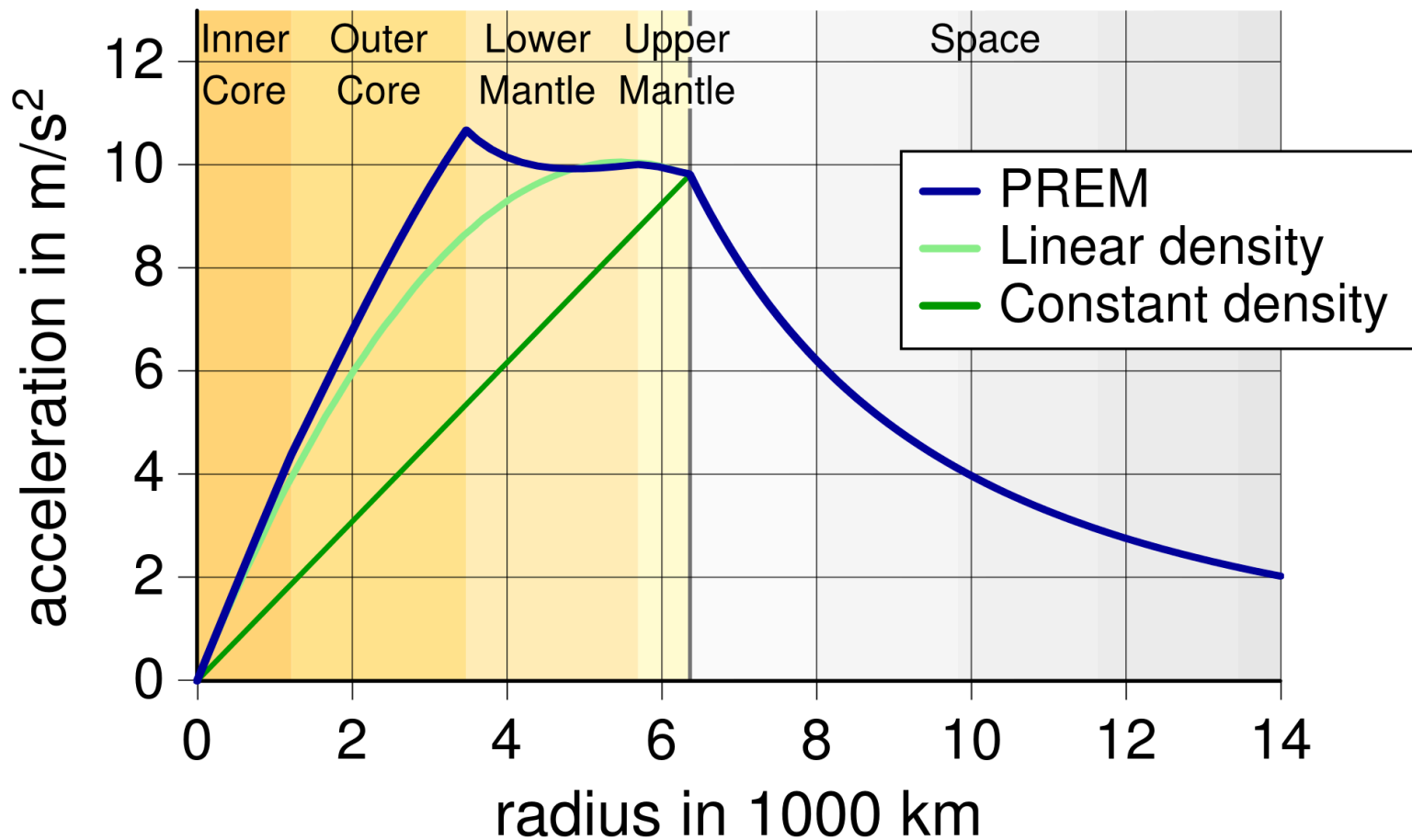
$$= -\frac{GM}{r^2} \hat{r} \quad r > R$$

⑨ $r = R$ these are equal

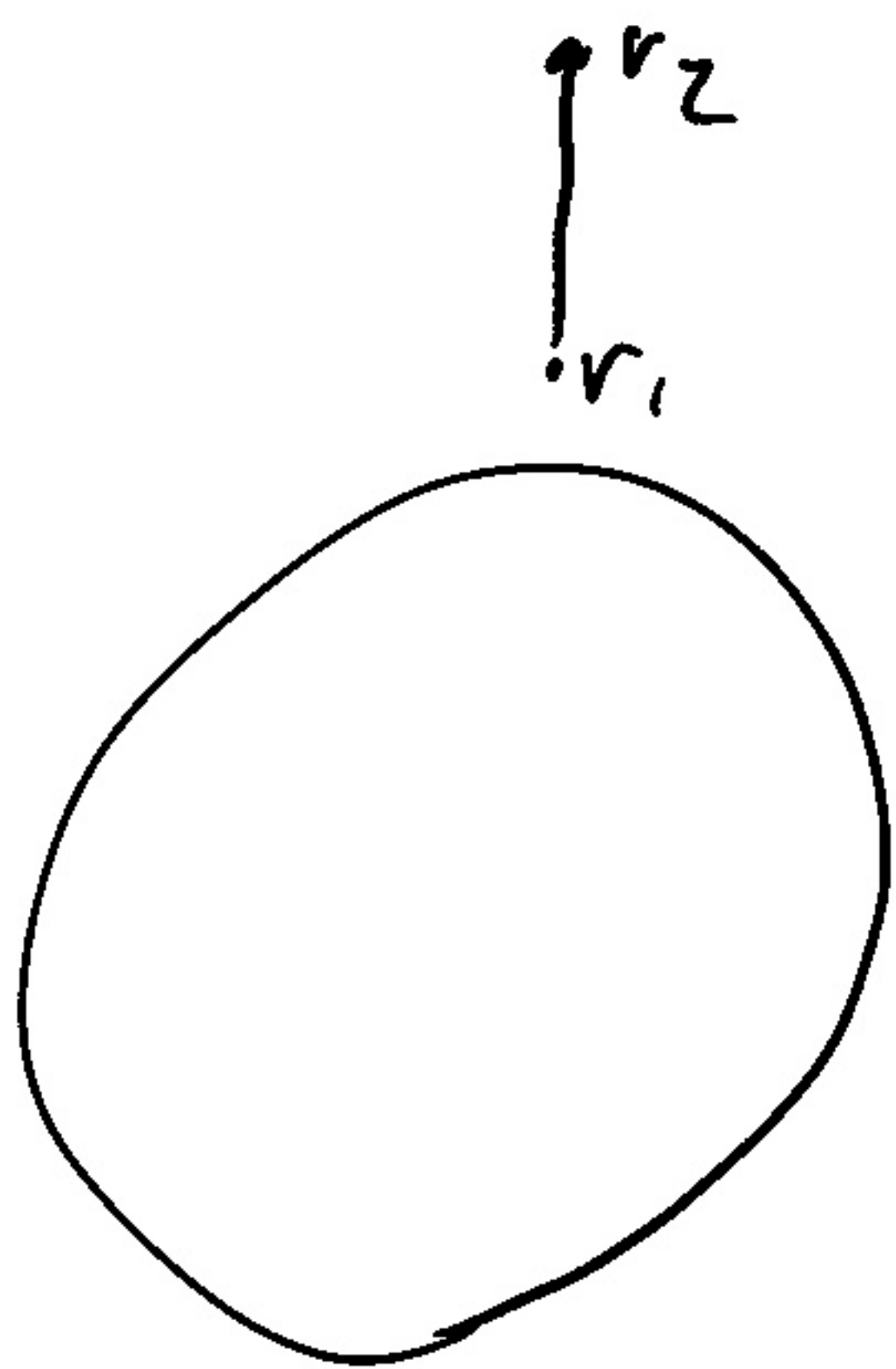


Actual Gravity Inside Earth

Free-fall acceleration of Earth



Gravitational Potential Energy:



$$W_{12} = \int_{r_1}^{r_2} \vec{F} \cdot d\vec{r}$$

- since force radial, only need to worry about r -coordinate

$$\vec{F} = -\frac{GMm}{r^2} \hat{r} \quad d\vec{r} = dr \hat{r}$$

$$\int_{r_1}^{r_2} \vec{F} \cdot d\vec{r} = \int_{r_1}^{r_2} -\frac{GMm}{r^2} \hat{r} \cdot dr \hat{r}$$

$$= \int_{r_1}^{r_2} -\frac{GMm}{r^2} dr$$

$$= \frac{GMm}{r} \Big|_{r_1}^{r_2} = \frac{GMm}{r_2} - \frac{GMm}{r_1}$$

$$\Delta U = -W$$

$$= \frac{GMm}{r_1} - \frac{GMm}{r_2}$$

$$= U_2 - U_1$$

convention is to set
 $U = 0$ @ $r = \infty$

say $r_1 = \infty$
 $U_1 = 0$

$$\Rightarrow \boxed{U = -\frac{GMm}{r}}$$

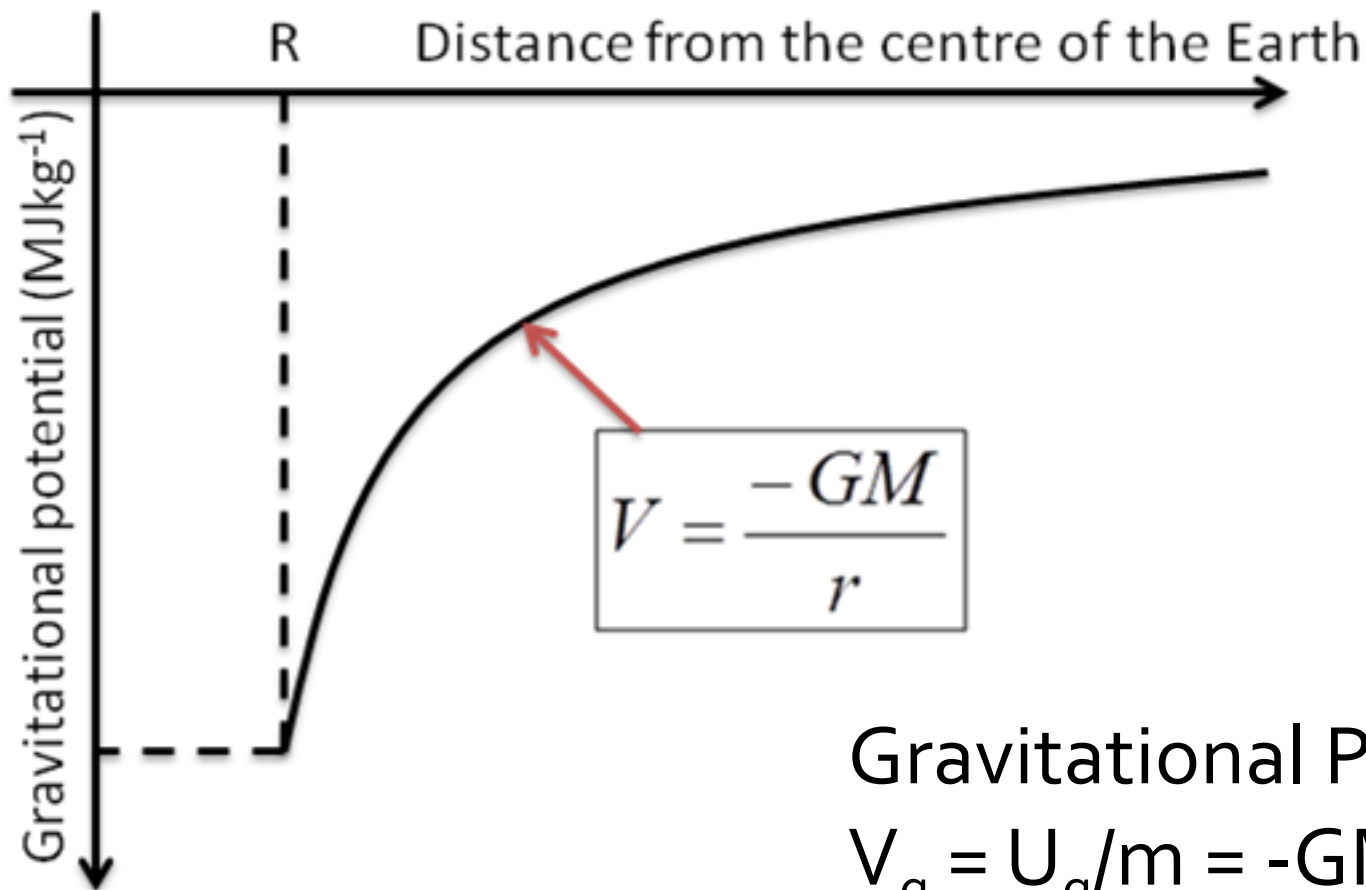
Note $F_r = -\frac{dU}{dr}$

$$= -\frac{GMm}{r^2}$$

= Newton's Law

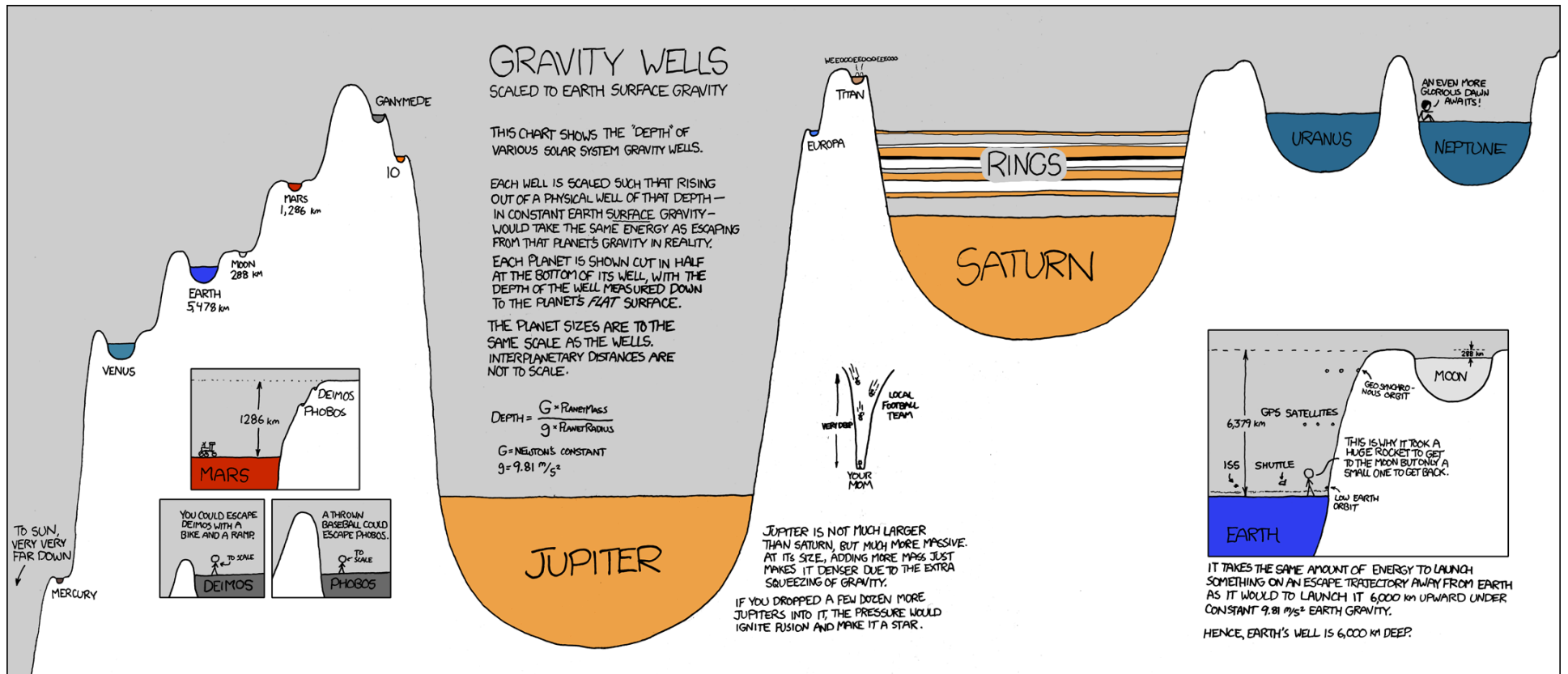
Gravitational Potential

- Depends on both masses involved, just like the force
- Normalize by mass to get a common quantity V_g



Gravitational Potential
 $V_g = U_g/m = -GM/r$

Obligatory XKCD



Concept Check

- To escape a body's gravitational pull, an object has to have non-negative kinetic energy at an infinite distance from the object. Knowing this, and the gravitational potential energy $-GMm/r$, can you predict the escape velocity?
- $\sqrt{(2GMm/r)}$
- $\sqrt{(2GM/r)}$
- $\sqrt{(GM/r)}$
- GM/r
- $2GM/r$